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# Enhancing VLBI capabilities: recent achievements and future upgrades of the INAF radio telescopes

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# Pinpointing **RE**peating **ChI**me **S**ources with **EVN** dishes, FRB 20201124A (Nimmo et al. 2022)

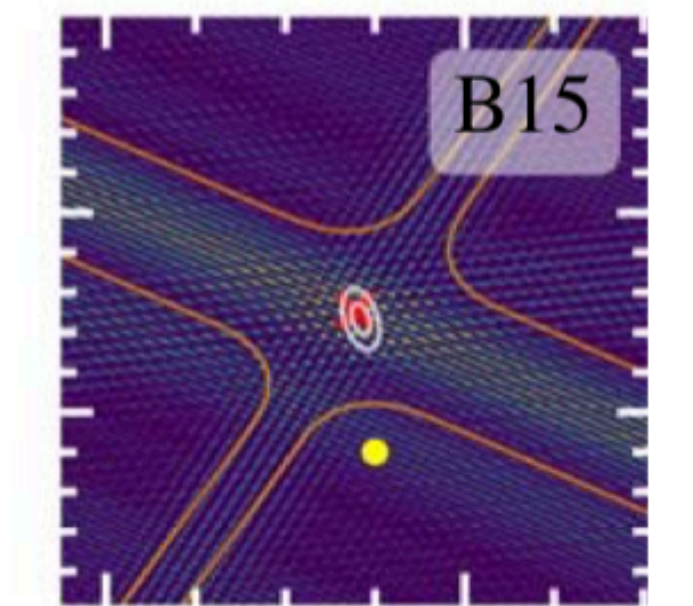
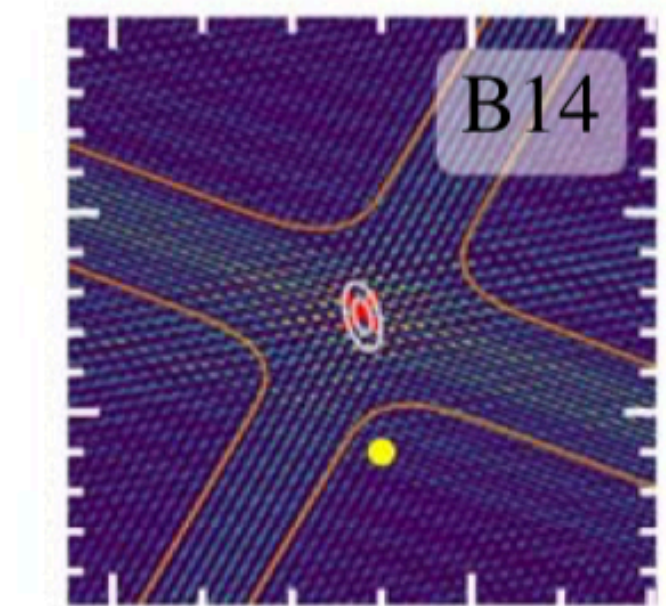
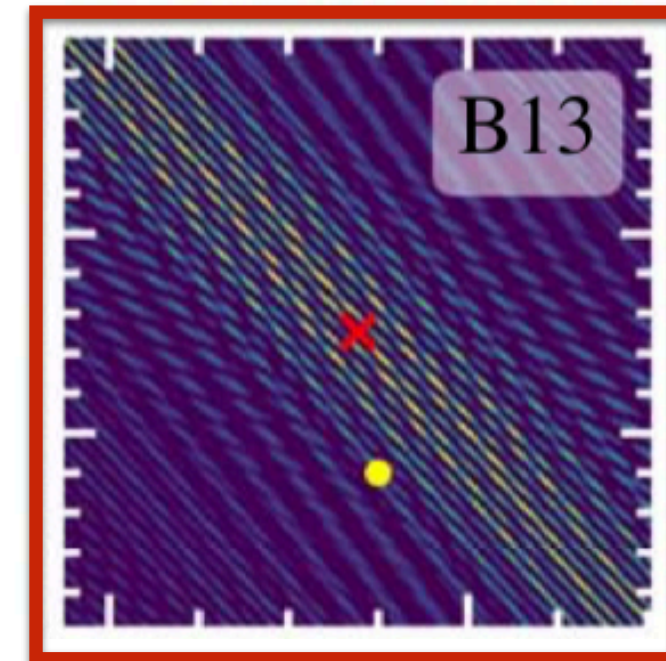
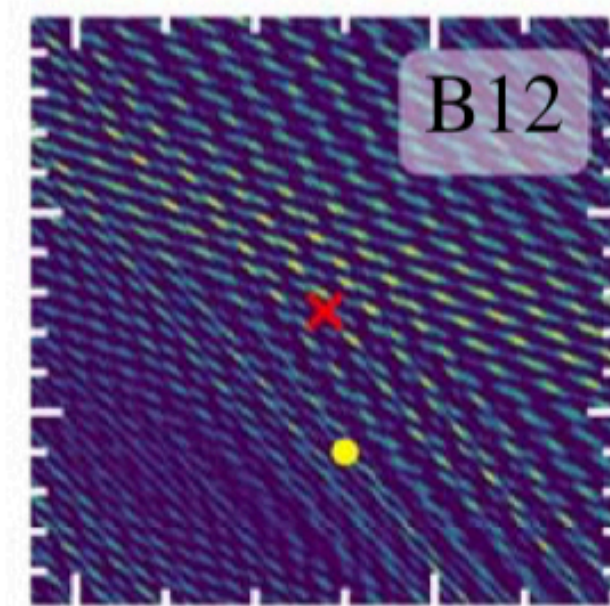
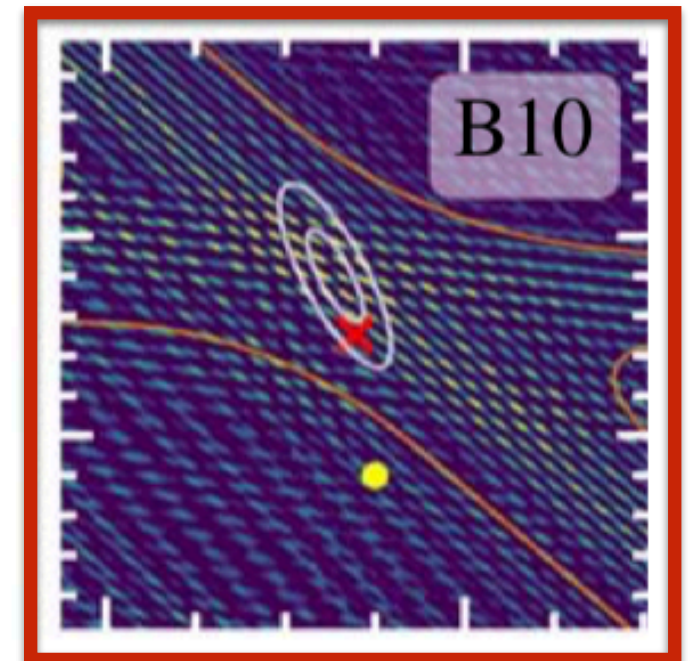
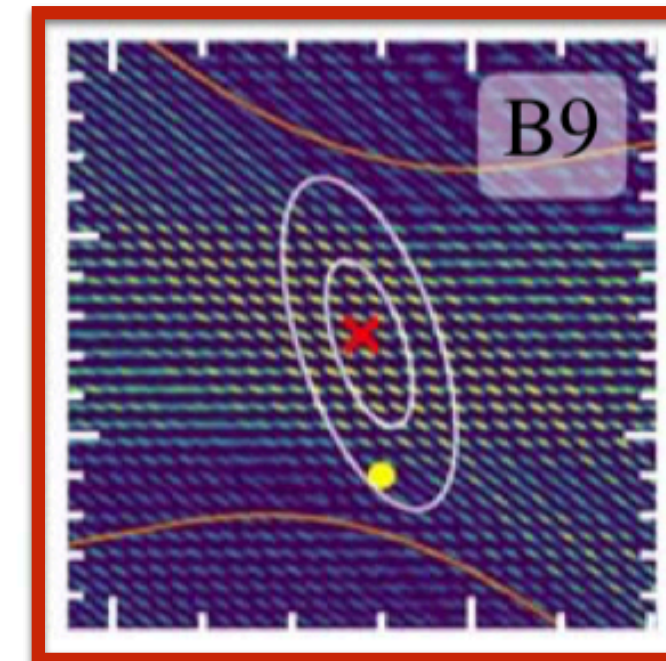
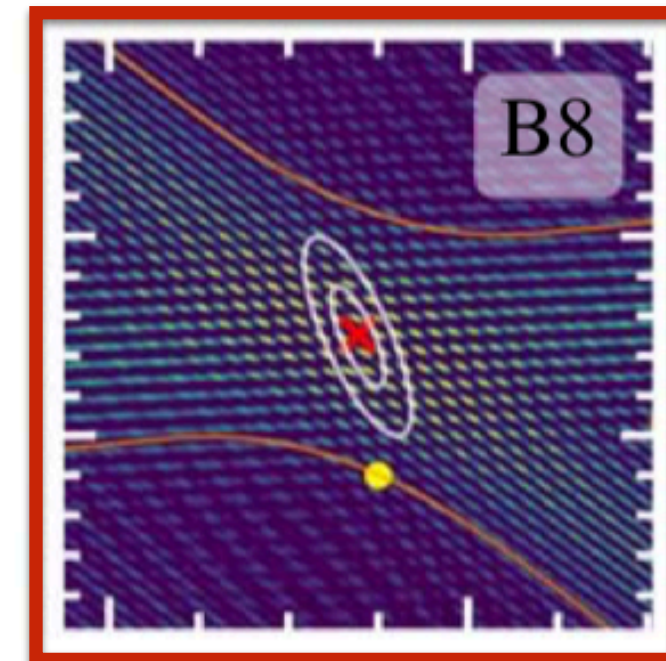
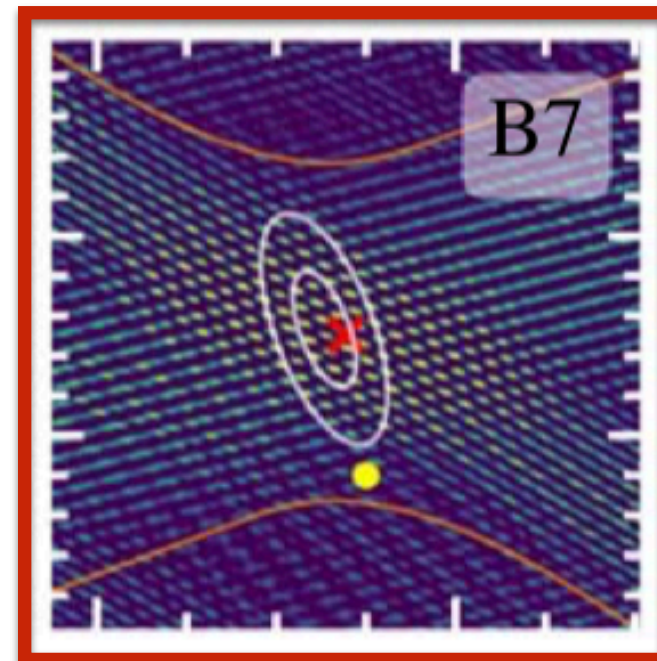
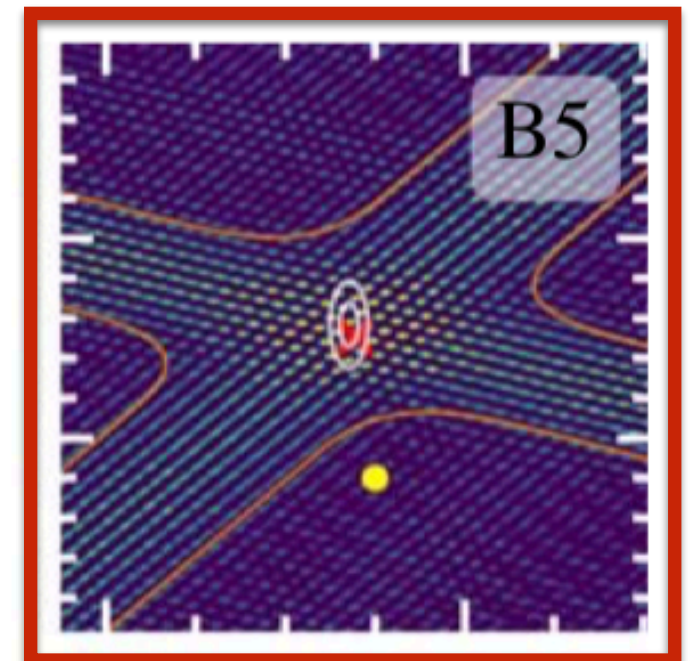
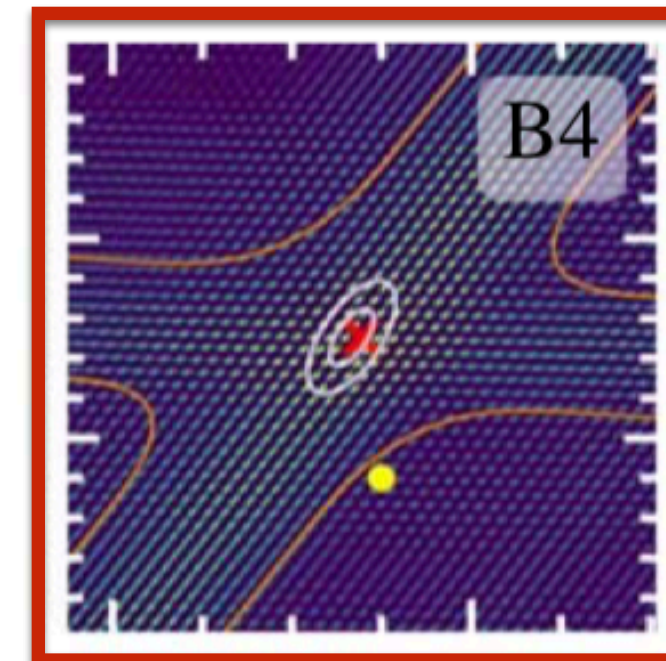
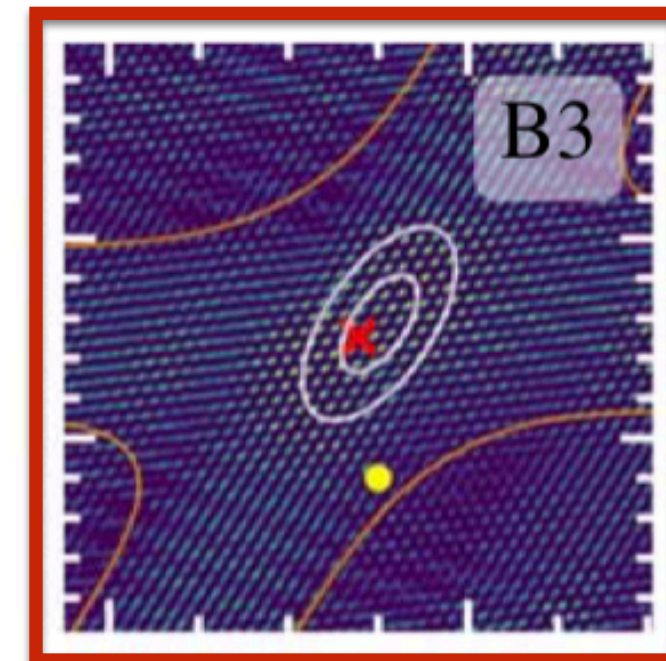
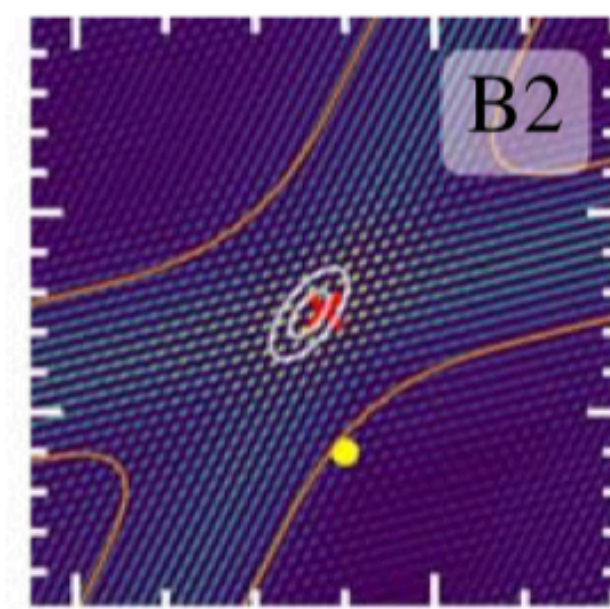


Mc, Nt, Sr

**Table 2**

Interferometric Burst Properties

Burst	Effective No. of Baselines <sup>h</sup>
Epoch 1	
B1	10
B2	10
B3	6
B4	6
B5	6
B6	15
B7	6
B8	6
B9	6
B10	6
B11	15
B12	15
B13	6



Sardinia, 64m



Medicina, 32m

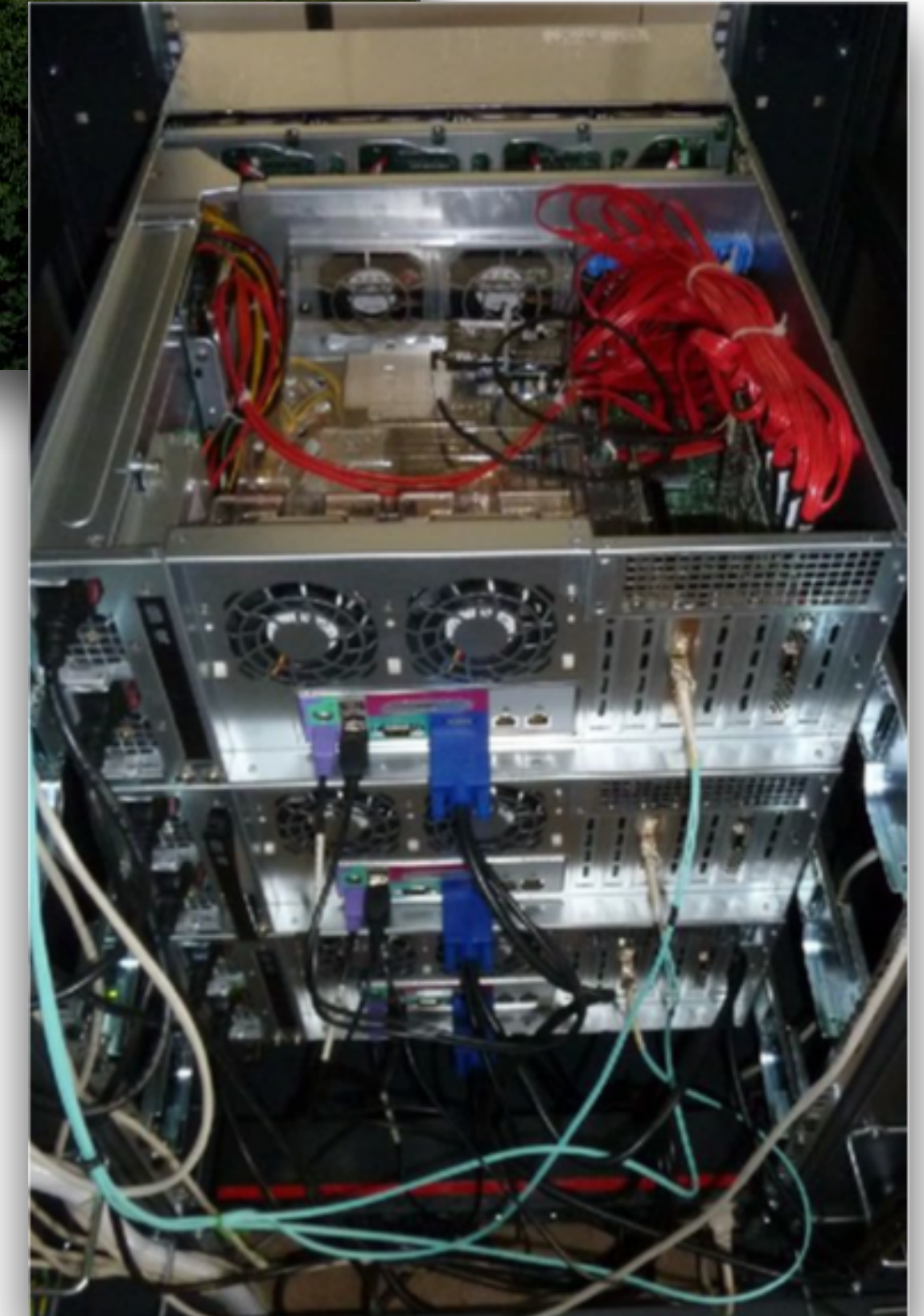


Noto, 32m

**INAF collecting area: 4825 m<sup>2</sup>**

**VLBA collecting area: 4909 m<sup>2</sup>**

Bologna,  
DiFX correlator



# INAF telescope tables



**Table 1.** INAF radio telescopes.

Location	Code	Size	Year	Latitude	Longitude	Elevation	Operating receivers
Medicina	Mc	32 m	1983	44° 31' 15'' N	11° 38' 49'' E	25 m	L, S, C, C+, X, K
Noto	Nt	32 m	1988	36° 52' 34'' N	14° 59' 21'' E	78 m	P, L, C+, K
Sardinia	Sr	64 m	2011	39° 29' 34'' N	9° 14' 42'' E	600 m	L, S, C, C+, X, K

**Table 2.** INAF VLBI baseline length and sensitivity.

Baseline	Length (km)	Sensitivity ( $\text{mJy min}^{-1} \text{ Gbps}^{-1}$ )		
		L-band	C+-band	K-band
Mc-Nt	893	5.9	7.8	6.1
Mc-Sr	592	1.8	1.7	2.5
Nt-Sr	580	1.8	1.9	2.7



# INAF radio telescopes - activities

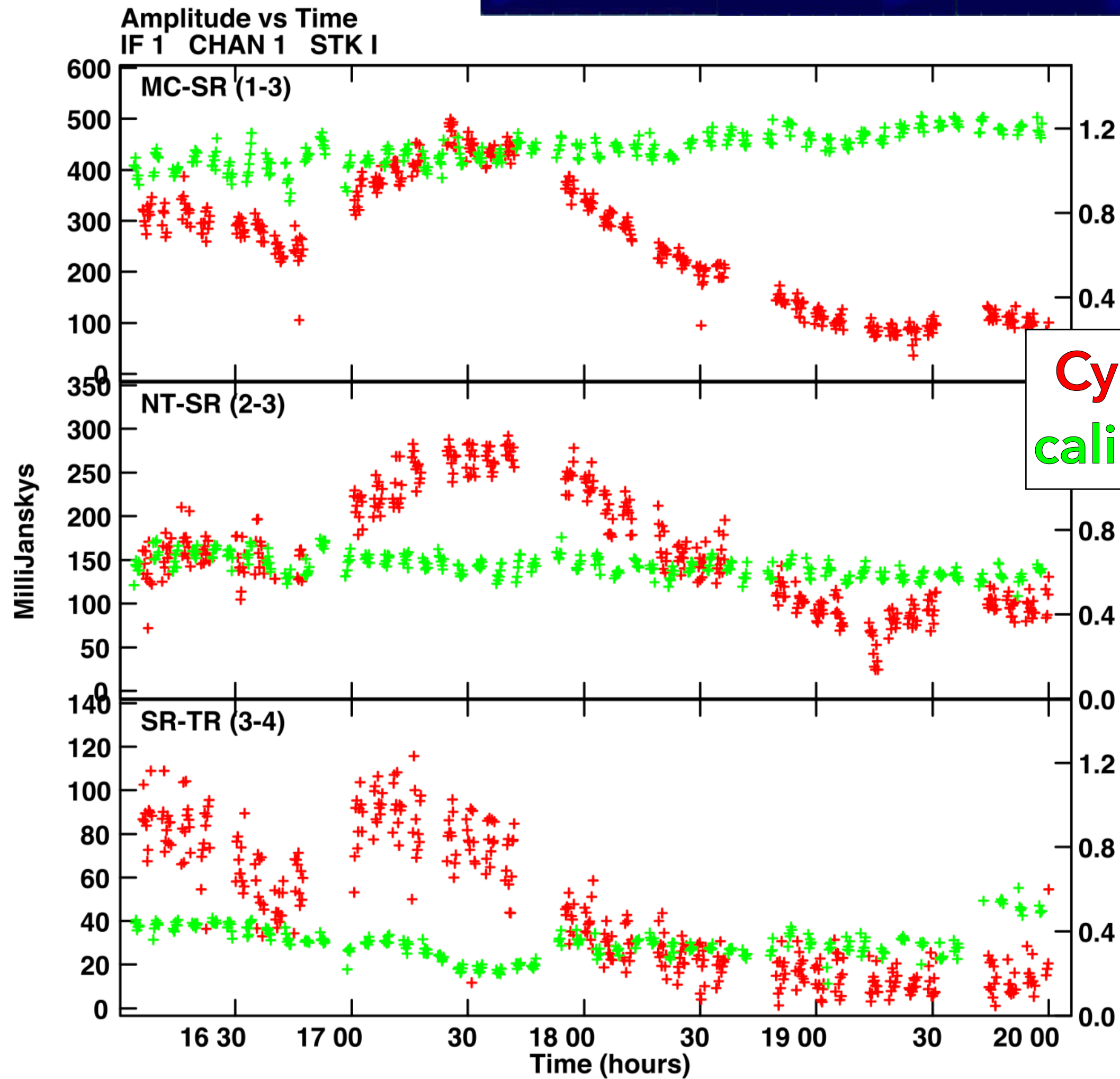
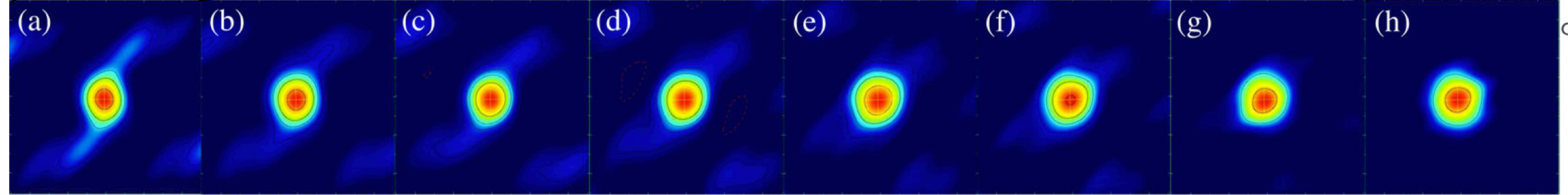


- ~70 days yr are devoted to EVN observations
- other commitments:
  - single dish observations
  - maintenance and development
  - IVS observations for geodesy (Mc, Nt), ASI support for space (Sardinia)
- other VLBI observations
  - RadioAstron, PRECISE (now EVN-lite)
  - East Asia To Italy: Nearly Global VLBI
  - PI-led, self-organised projects

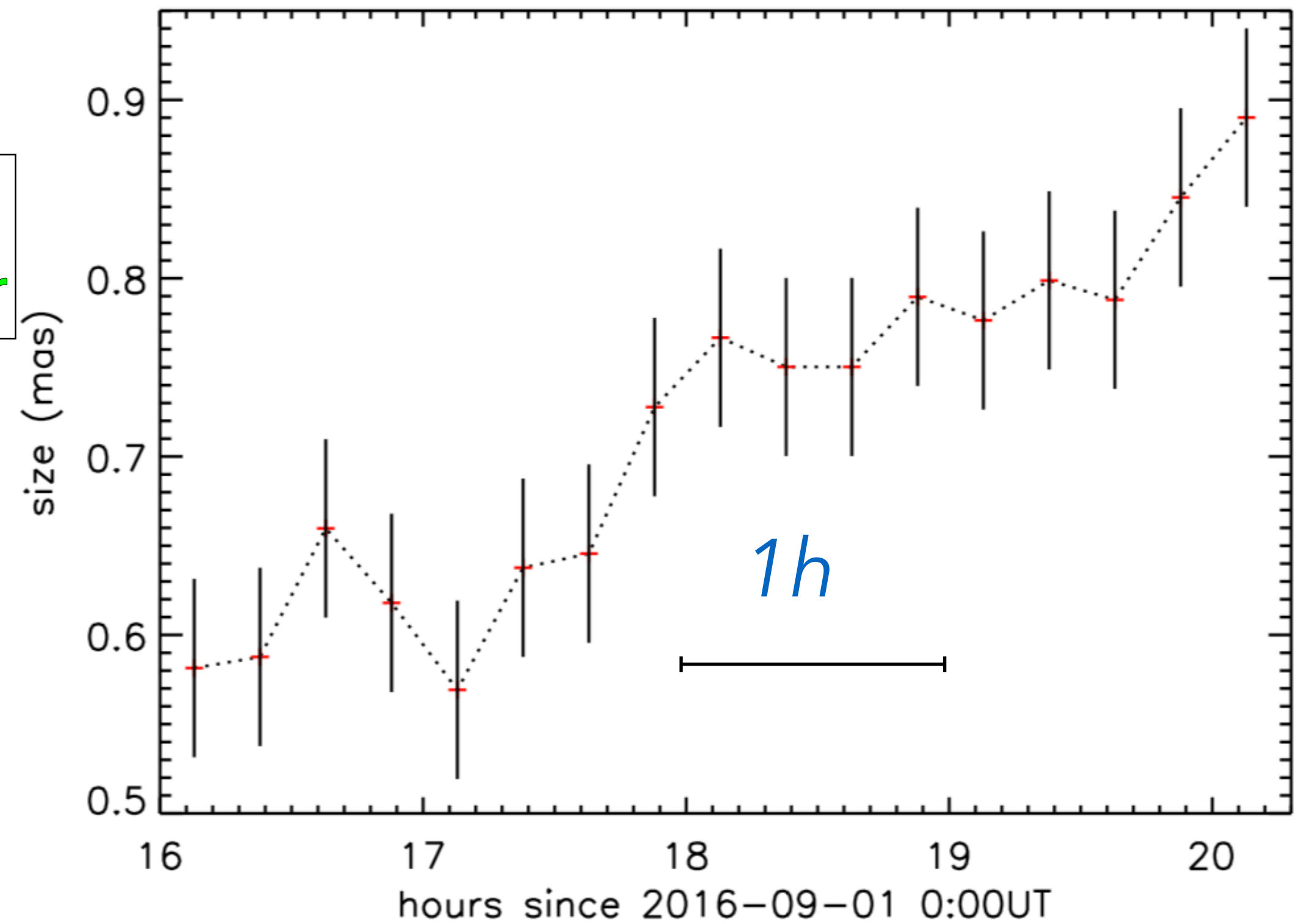


# Cyg X-3

*Egron et al. 2017, first Italian VLBI paper (+Ys/On/Tr)*

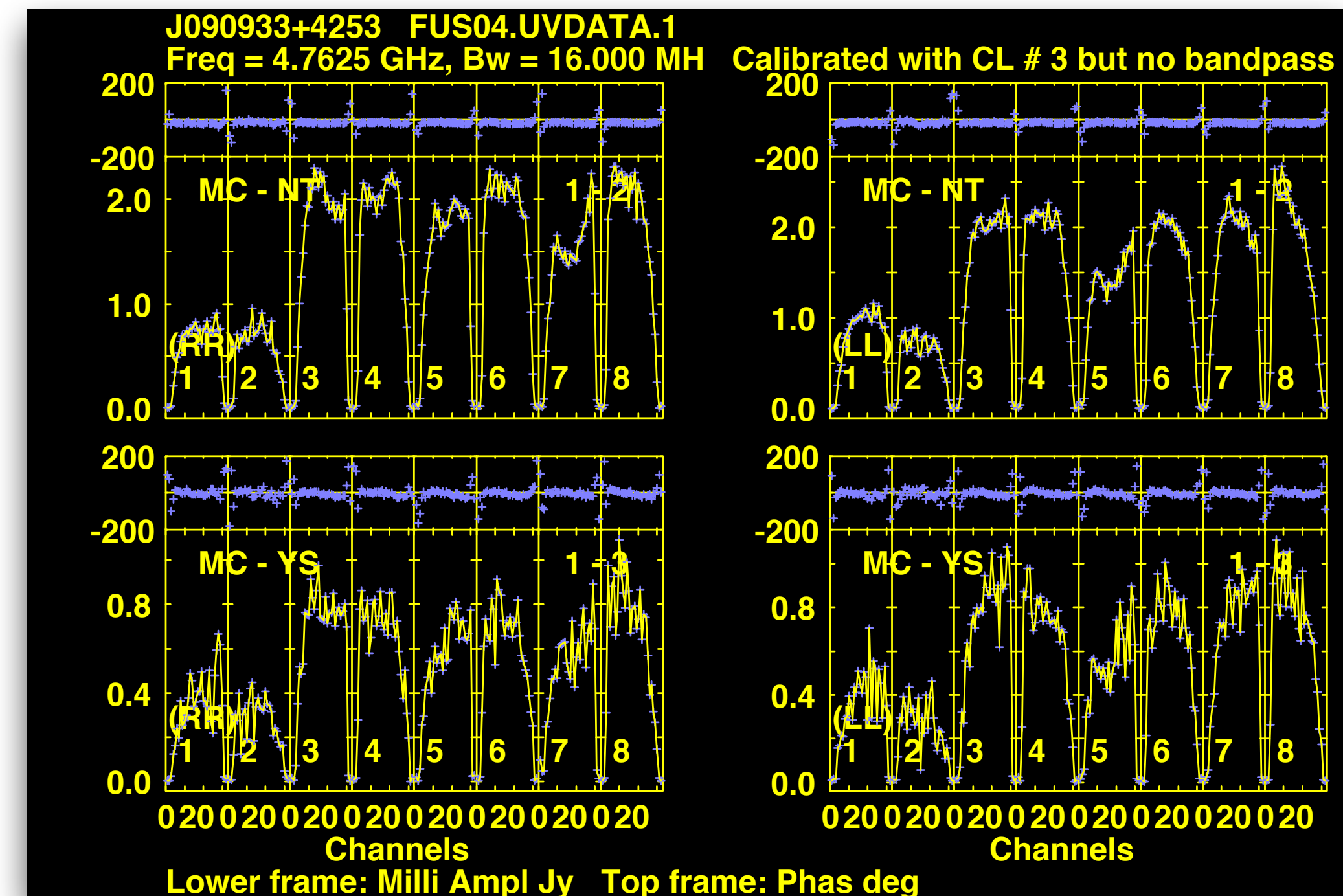
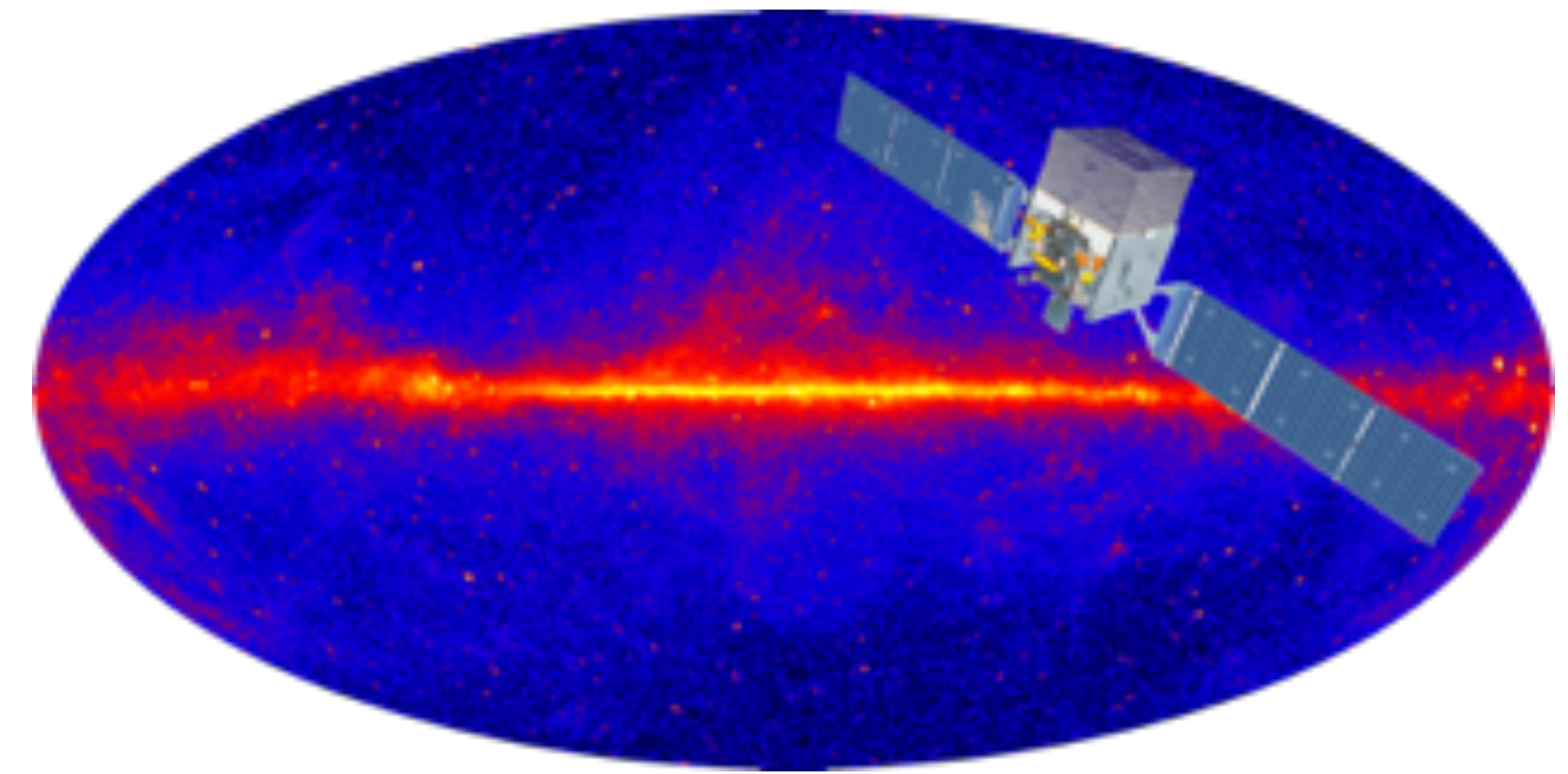


**Cyg X-3**  
**calibrator**



# VLBI observations of $\gamma$ -ray unidentified sources

- ~30% of sources in Fermi-LAT catalogues have no clear counterparts ( $\sim 10^3$  sources north of -20deg)
- some empty fields, some just too many arcsec-scale candidates (especially at low fluxes)
- filter out the intruders with VLBI!
- potentially on massive scales
- 5 GHz, baseline sensitivity of 10 mJy, 3-5 mas astrometry
- stations: Mc, Nt, Ys, Tr, Ir



# Upgrades

1. High-frequency: compact triple band receivers on Mc/Nt/Sr
2. Mid-frequency: broadband receivers on Mc/Nt

# Upgrade #1

- “Enhancement of the SRT for the study of the Universe at high radio frequencies” (scientific coordinator Federica Govoni, ~20 M€)
- 9 WPs, including WP4 “Simultaneous microwave compact triple-band (**CTR**) receiving system for the three Italian radio telescopes (18-26; 35-50; 85-116 GHz)”
- Receiver specifications by INAF, design and construction by KASI
- CTR’s delivered and currently being installed on INAF telescopes

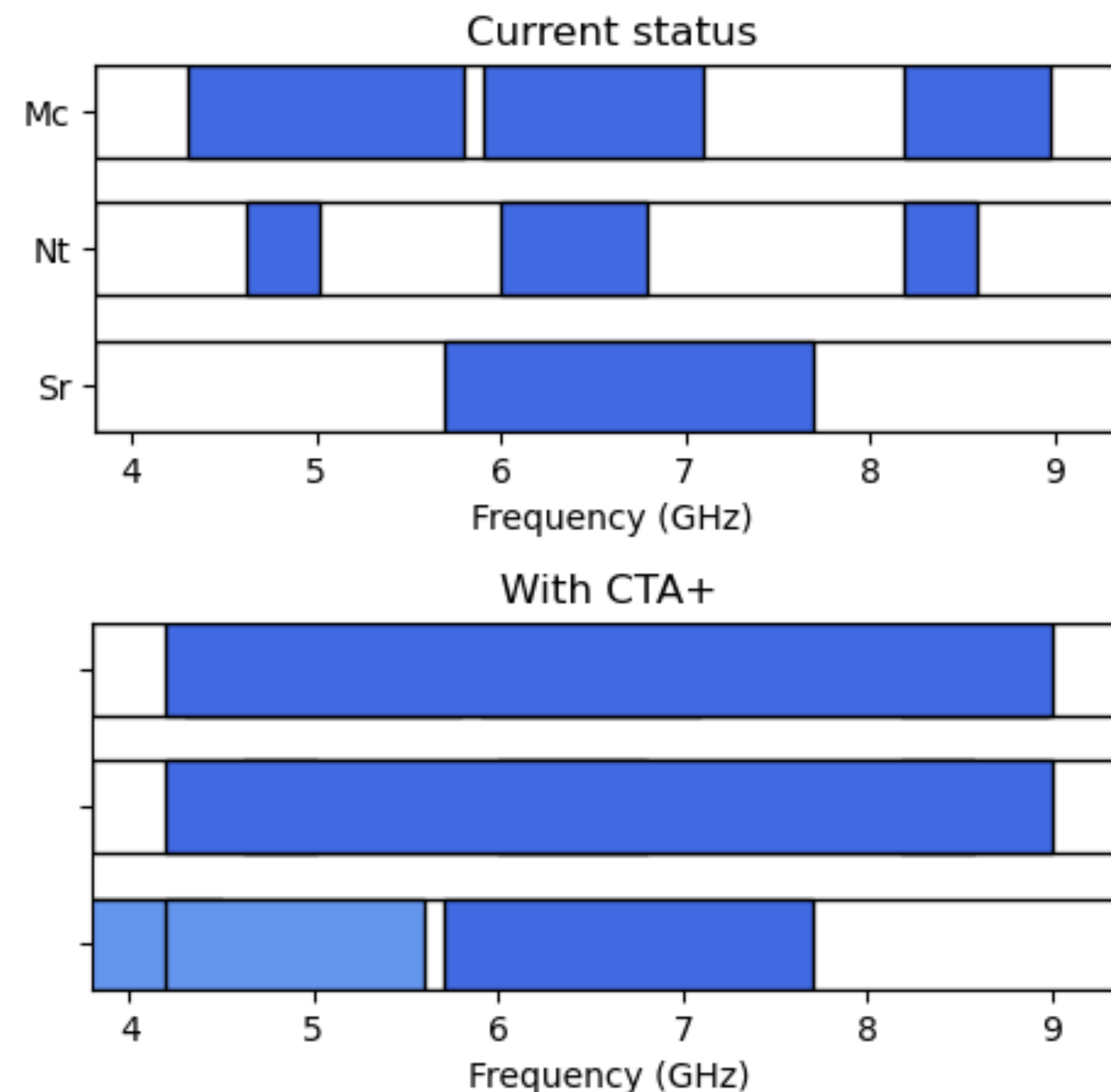


# Current status

- Sardinia: CTR in position on the telescope, initial tests carried out (8/2024)
- Medicina: installation of active surface system starting this month (9/2024) and new sub-reflector surface
- Noto: active surface available but new sub-reflector required - several major upgrades on the antenna ongoing

# Further upgrades: "CTA+"

- 70M€ INAF-led project to improve Cherenkov Telescope Array Observatory
  - and supporting MWL instrumentation
- VLBI key for transient physics
- Complements "PON" at lower frequency:  
~1M€ for Medicina&Noto broadband (4-9 GHz) receivers
- Delivery: end 2025





# Outlook

- Multimessenger context
  - Gravitational wave astrophysics (Ligo-Virgo, Einstein Telescope): need screening of candidates to select primary targets to follow up with full-sized VLBI arrays
  - Neutrino astrophysics (IceCube, KM3NeT): large number of potential source types/counterparts, monitoring projects
- compact objects & other transients (VRO-LSST; CTAO; etc);  
gravitational lenses

***...stay tuned!***