

Latest results

from



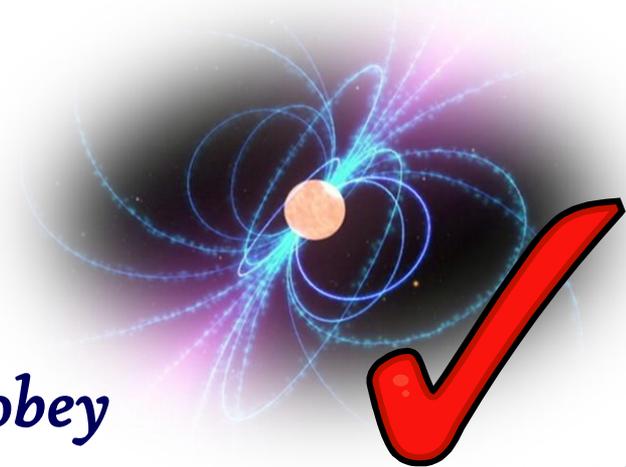
LOFAR observations

of

polarised pulsars



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Latest pulsar polarisation results with LOFAR

Thursday 24th November 2011

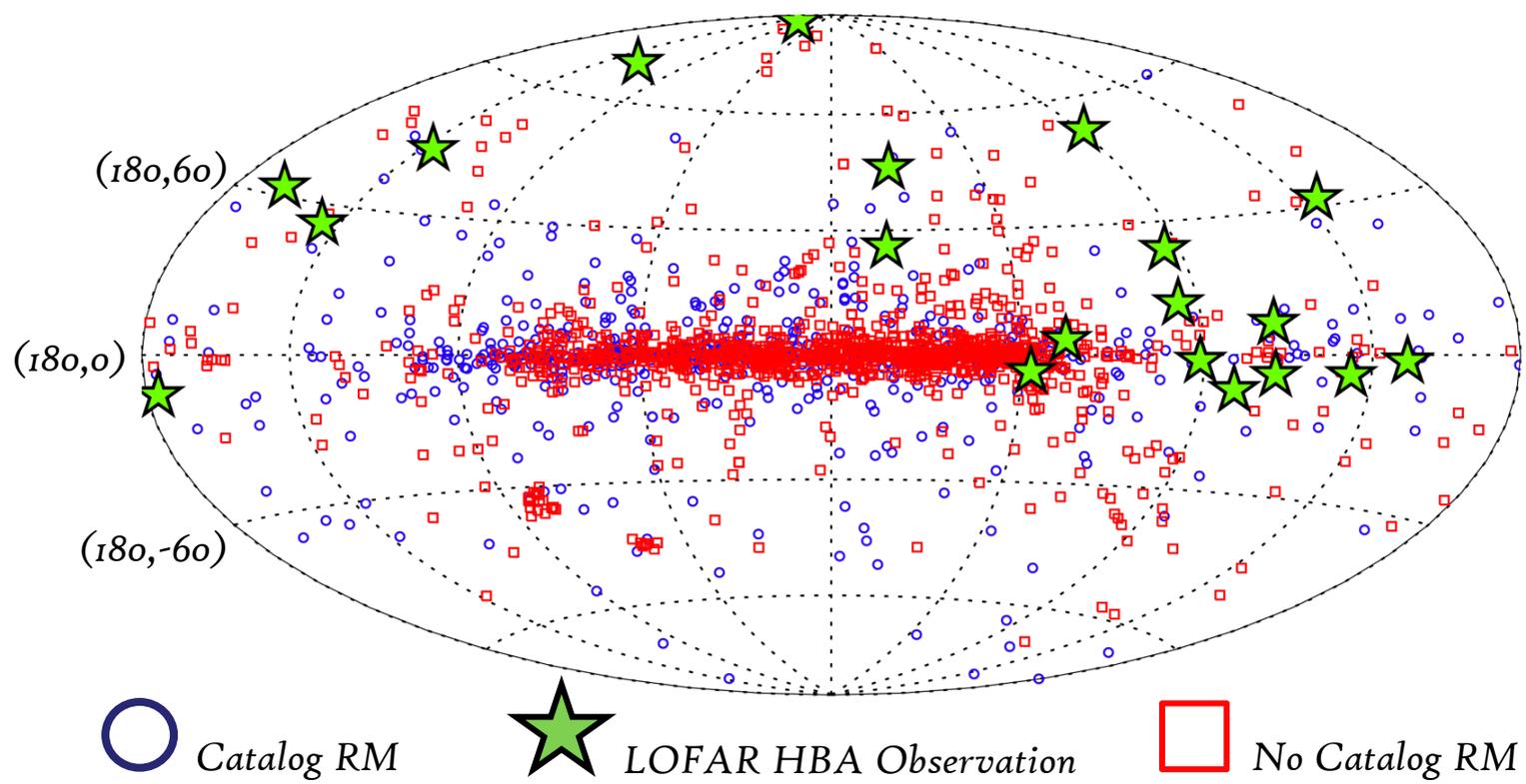
Outline

- ⊙ New(ish) observations and results from:...
 - ⊙ RM synthesis on polarised pulsars
 - ⊙ RM tracks of Ionospheric variations
 - ⊙ Almost-simultaneous beam-formed/imaging obs
 - ⊙ Giant pulses from the Crab pulsar
- ⊙ Future prospects for observations etc...

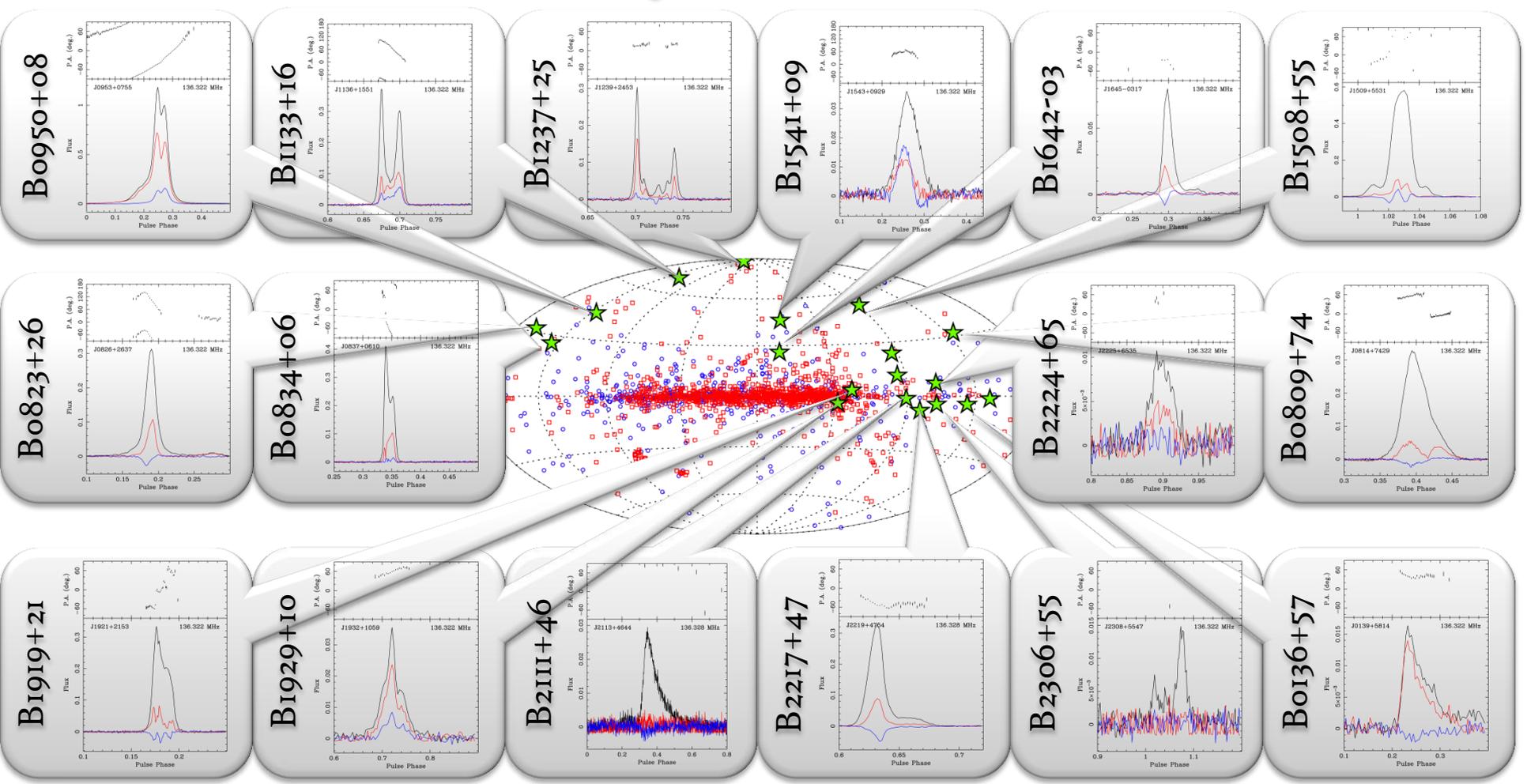
In collaboration with the MKSP & Pulsar Working Group (TKSP)

Previously... March 2011

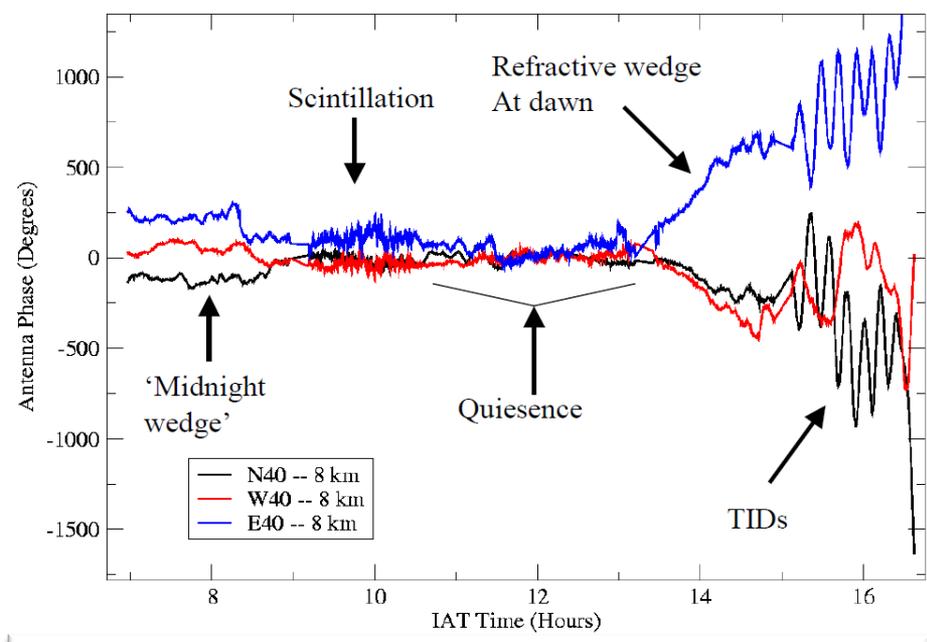
10 mins, Superterp stations, coherently summed, raw voltages
 Freq = 136.322 MHz BW = 6.25 MHz



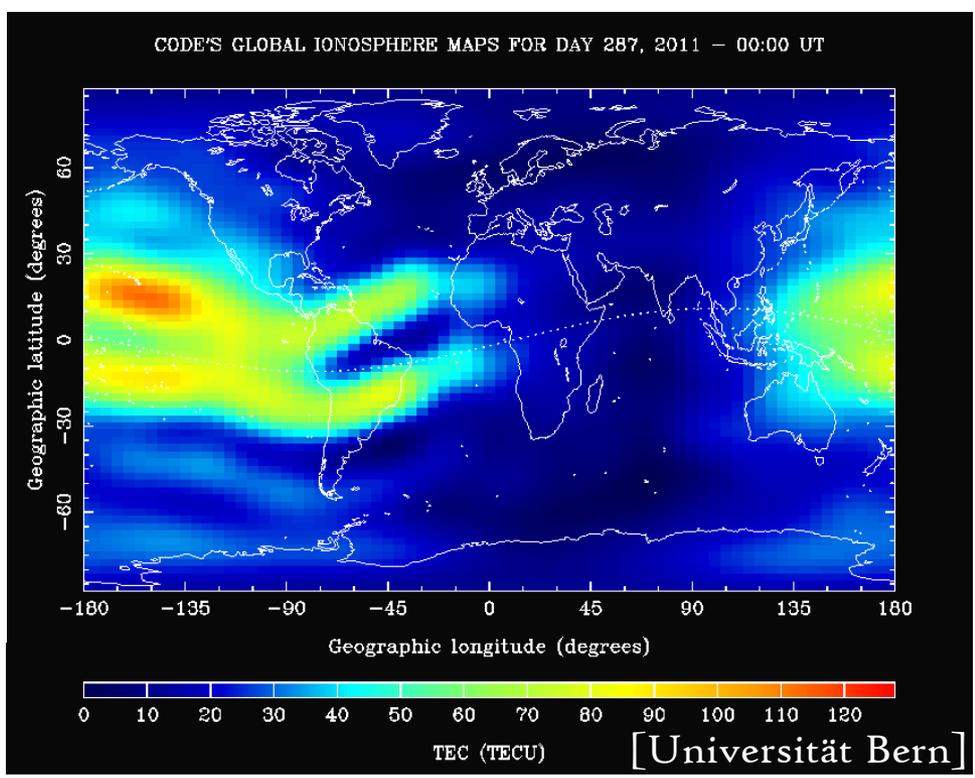
Previously... March 2011



So... The Ionosphere...



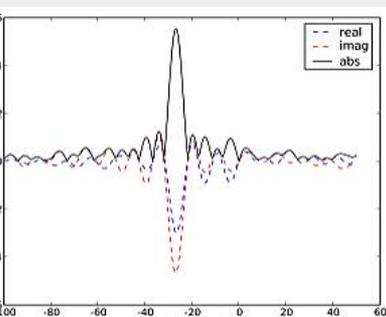
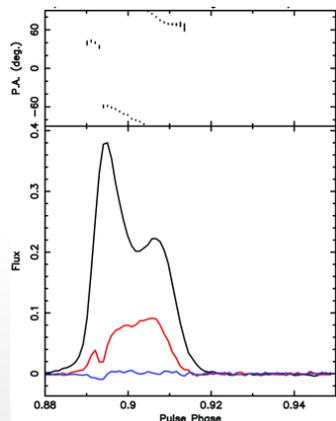
Relative antenna phase over 8 hr VLA obs
Vir A at 78 MHz
Kassim et al. 2007



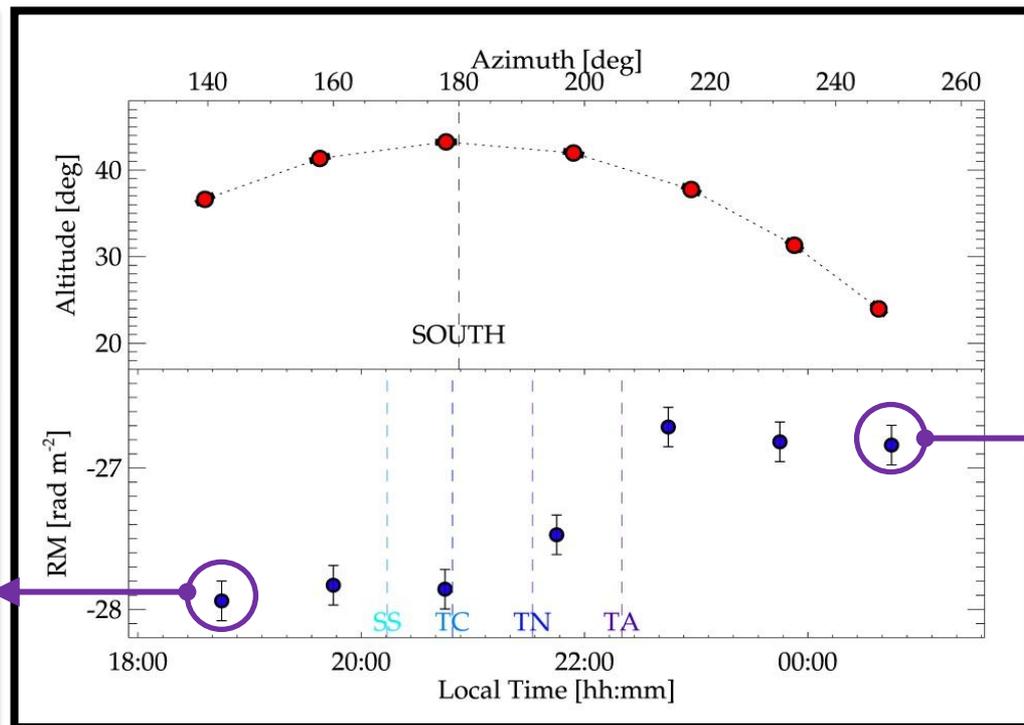
An Ionospheric Journey...

- ⊙ First observations: 11th April 2011 (sunset)
 - ⊙ 10 minute observations every hour over 6 hrs incl.
 - ⊙ Superterp raw voltage Freq = 122.65 MHz BW = 6.25 MHz
- ⊙ Further investigation: 20th October 2011 (sunrise)
 - ⊙ 3 minute observations every 10 minutes over 3 hrs incl.
 - ⊙ Superterp raw voltage Freq = 134.18 MHz BW = 9.155 MHz

Initial results at sunset...



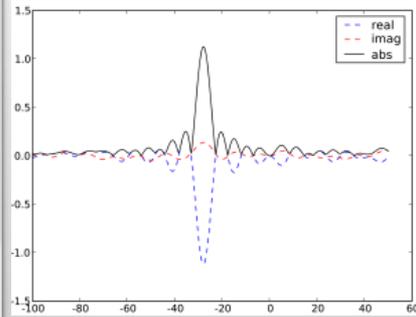
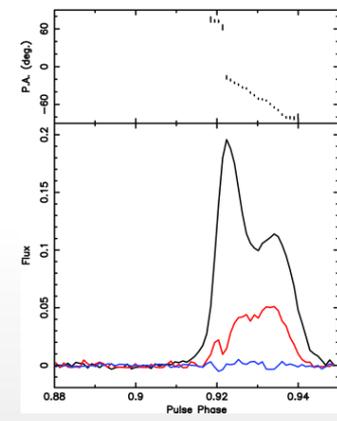
RM = -27.94 ± 0.15



PSR Bo834+06

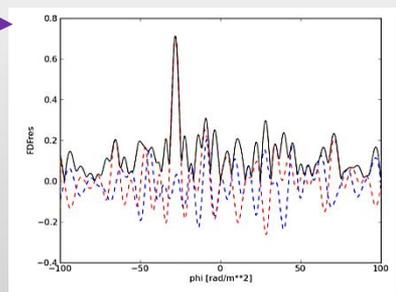
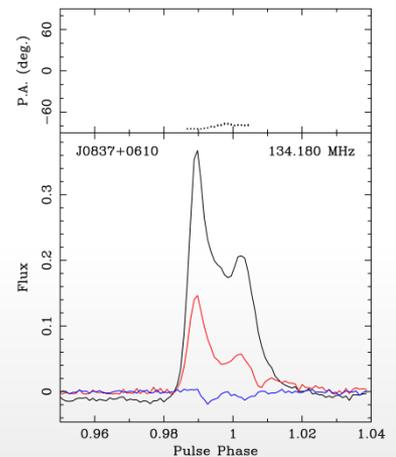
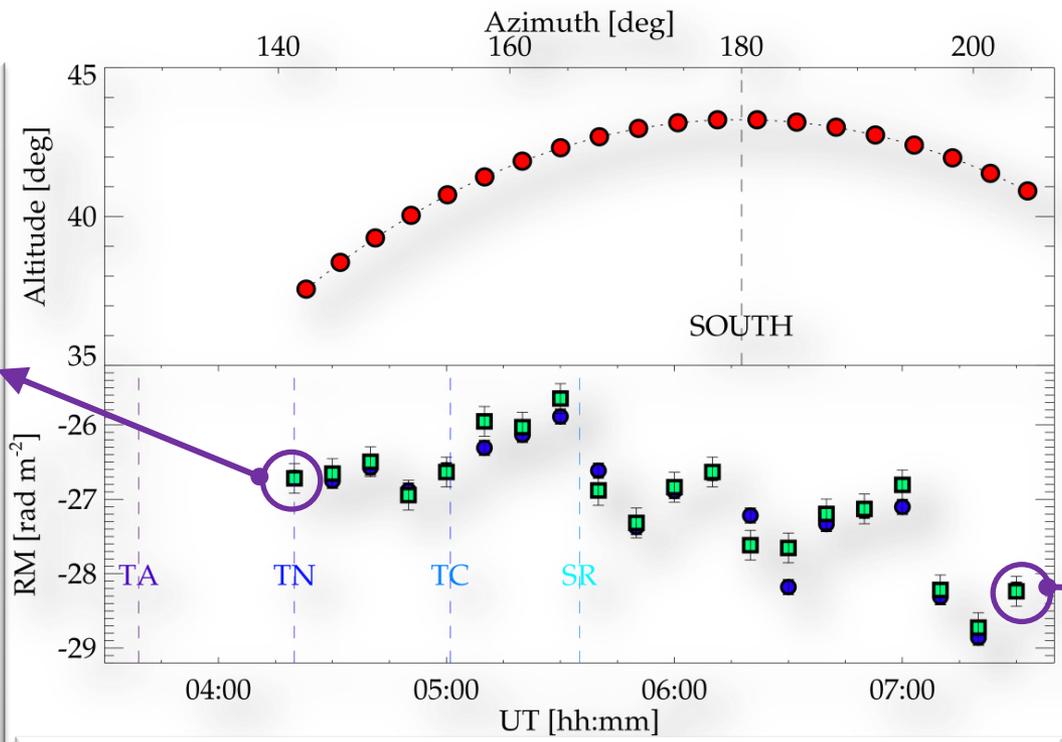
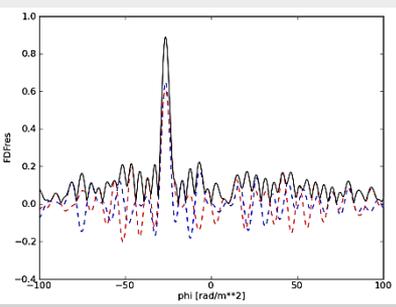
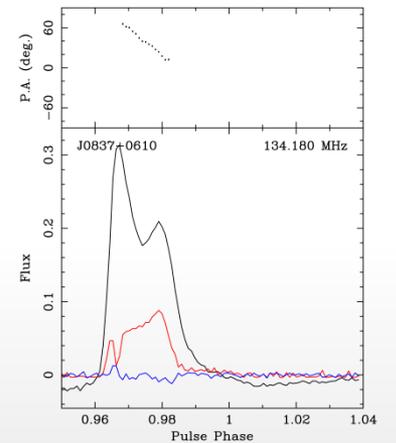
SS = Sunset

T[C,N,A] = Twilight [Civil, Nautical, Astronomical]



RM = -26.84 ± 0.07

Further results at sunrise...



SR = Sunrise

T[C,N,A] = Twilight [Civil, Nautical, Astronomical]

N.B.: For true RM of pulsar multiply by -1

$$RM = -26.7 \pm 0.2$$

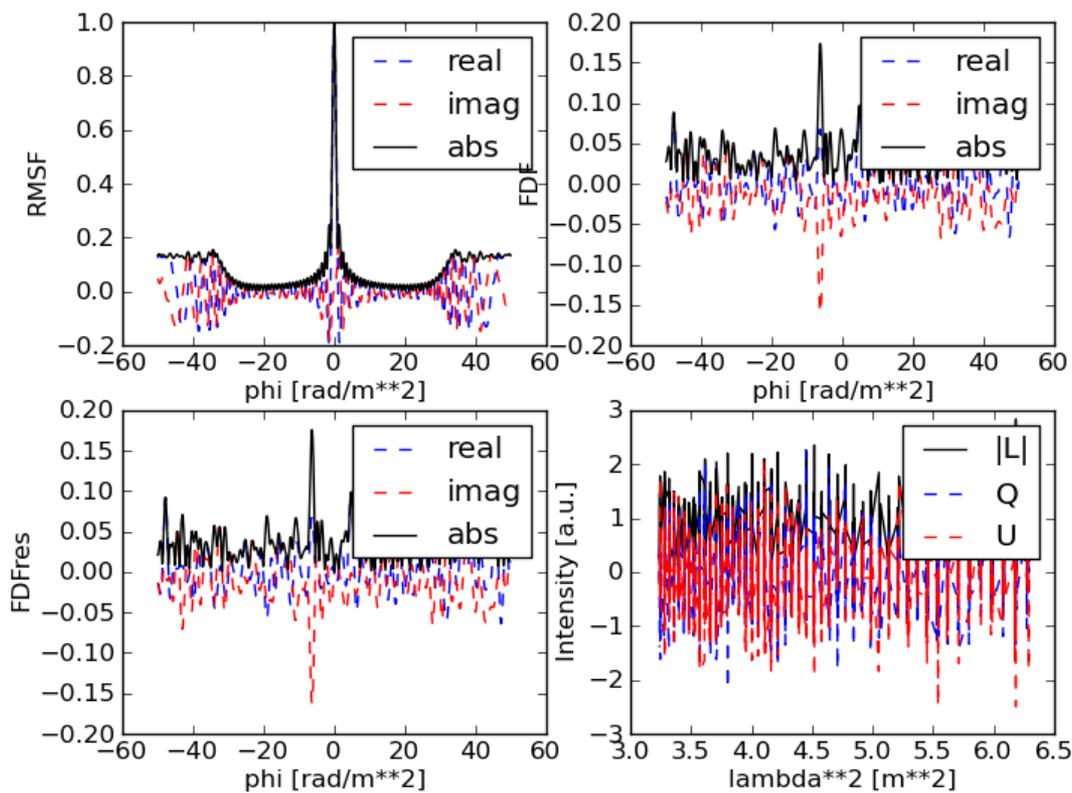
$$RM = -28.2 \pm 0.2$$

Another attempt... with imaging!

- ⊙ New observation with 'simultaneous' BF and IM modes:
 - ⊙ 3 mins beam-formed tied array with Superterp PSR B0823+26
 - ⊙ New aspect: non-continuous sub-band numbers
 - ⊙ 15 mins imaging mode with 19 CS & 9 RS PSR B0823+26
 - ⊙ 1 min calibrator field 3C 123
 - ⊙ See Mike Bell's presentation?!?!
 - ⊙ 1 min gaps between each...
- ⊙ BF: 13th – 14th November 2011, 3 min obs every 19 mins over 9 hrs

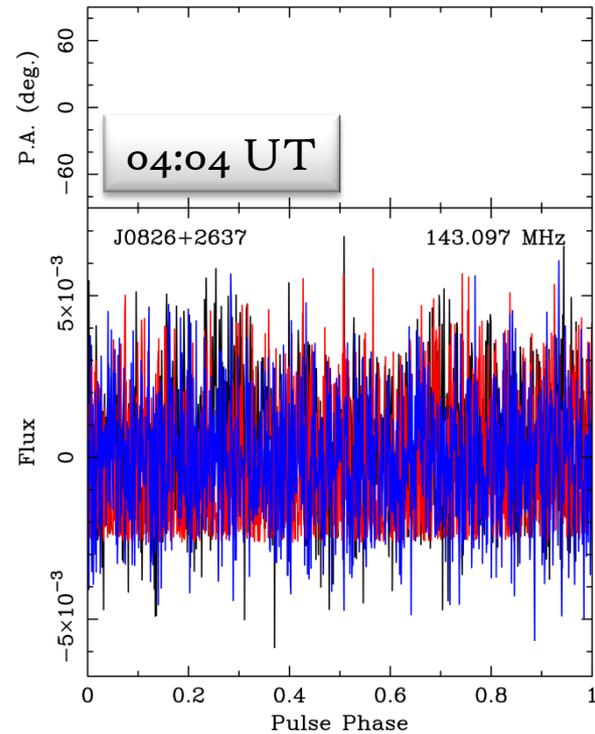
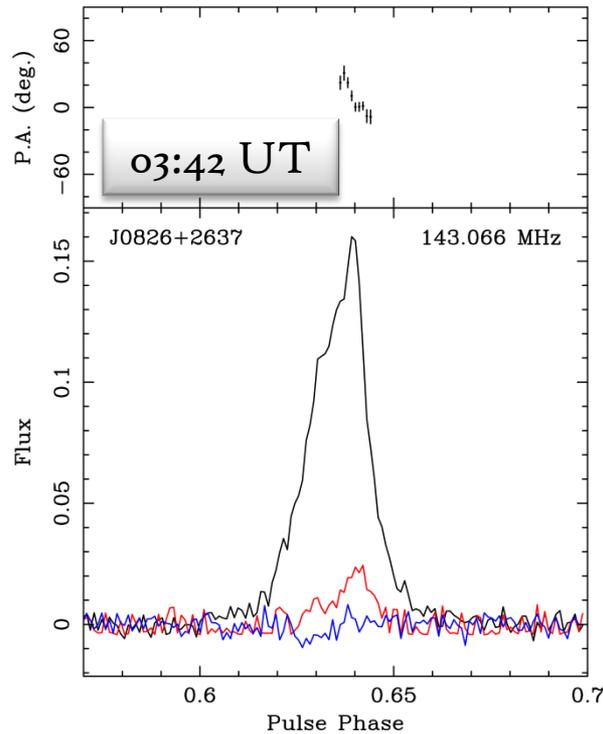
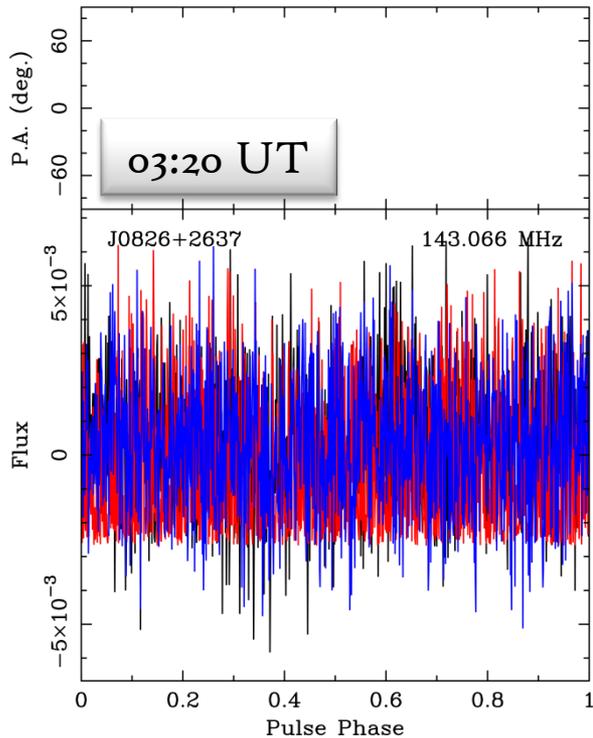
Good news...

© RM synthesis FWHM = 1.25 rad m^{-2} ... best to date...
 < 1.5 min integration, $\text{SNR}_{|L|} \sim 20$, $\text{RM} = -6.17 \pm 0.06 \text{ rad m}^{-2}$



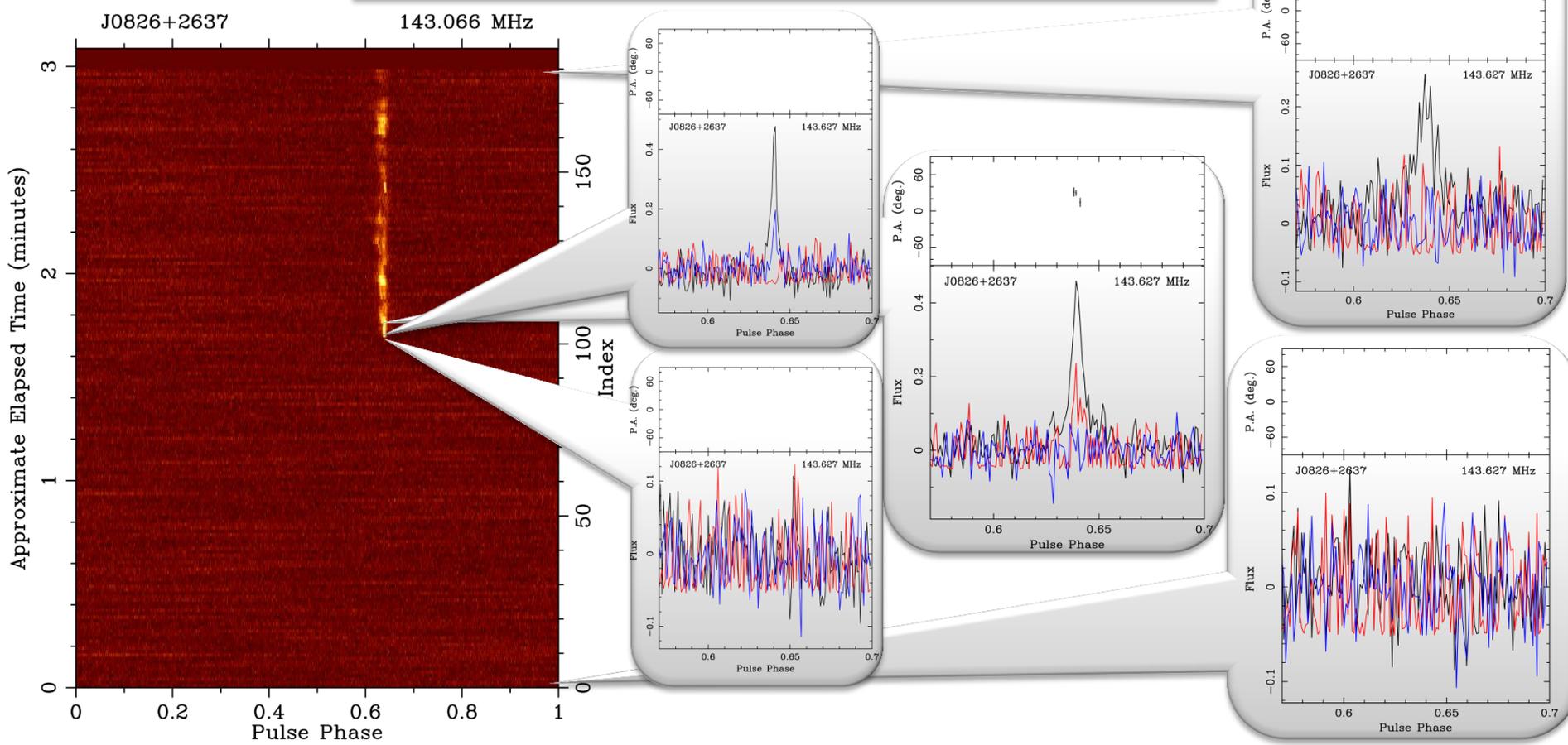
'Bad' news...

- ⊙ It turns out that B0823+26 is a nulling pulsar.... Oooops!
- ⊙ Not the best for tracking RM or calibration... BUT



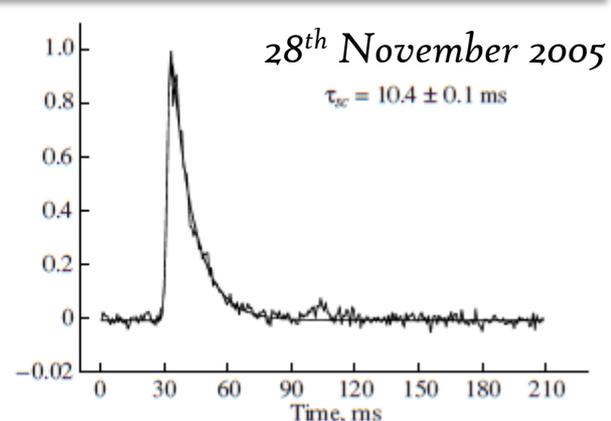
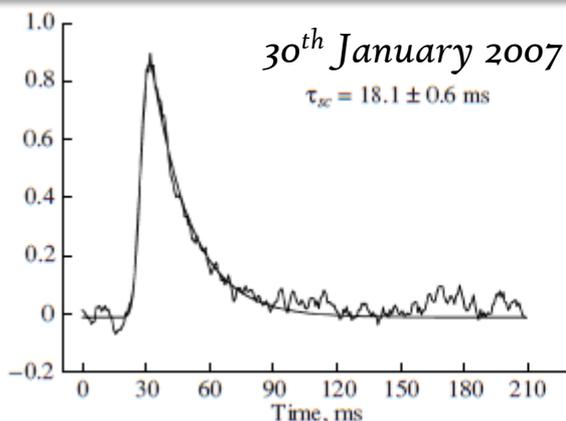
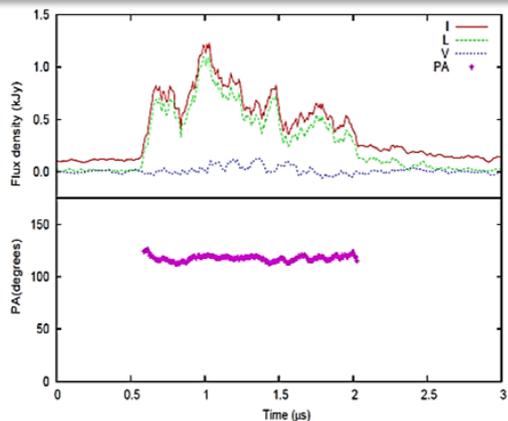
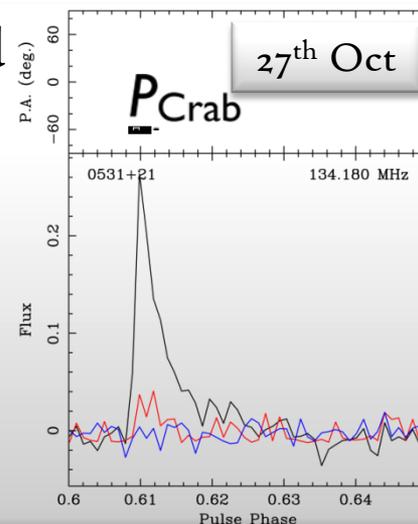
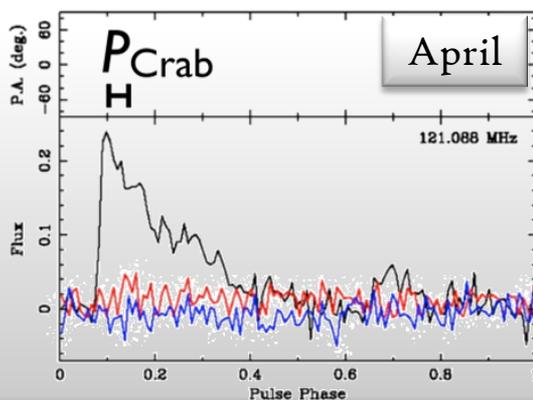
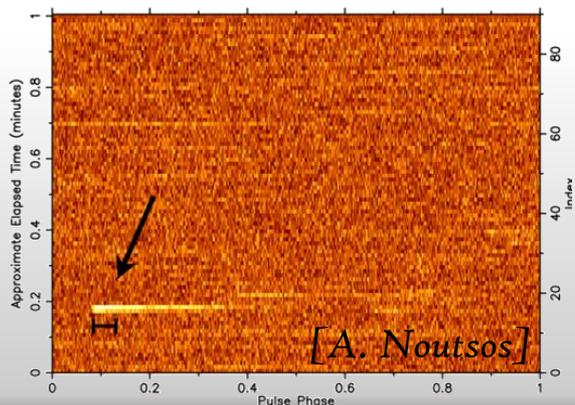
On the bright side...

Very interesting in other ways...



Crab giant pulses...

Work in progress... 30 min obs, 27 detected



15.1 GHz at Effelsberg [Jessner et al. 2010]

111.88 MHz at LPA [Smirnova & Logvinenko 2008]

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The future looks bright...



- ⊙ Past obs have proven interesting in many ways...
 - ⊙ RM can be achieved with high precision - Ionosphere
 - ⊙ Looking at consistency with imaging data with view to calib
 - ⊙ Plus some interesting things on the way: 0823 and the Crab
- ⊙ Near -future observations:
 - ⊙ Repeat of BF-IM obs with 0834 (not nulling - as we know!)
 - ⊙ More ionospheric studies... i.e. > 4 pulsars over various l.o.s.

$$E = mc^2$$



Thank you for listening!