

Latest commissioning results of GRG

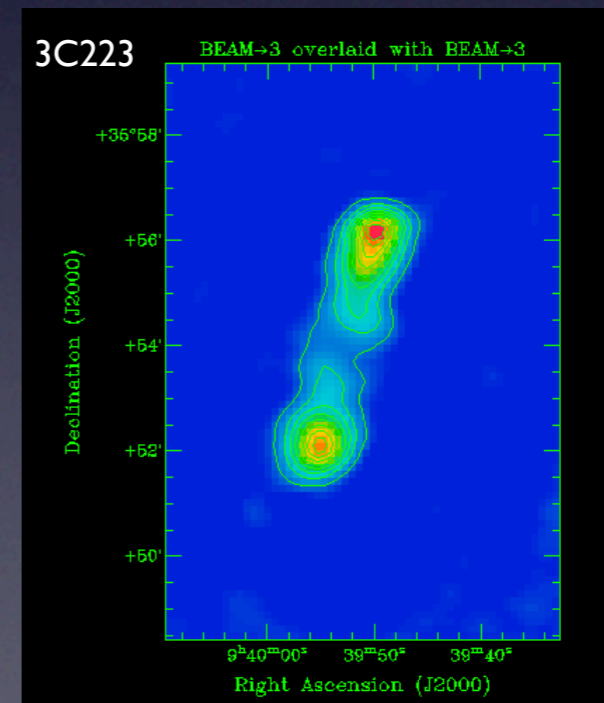
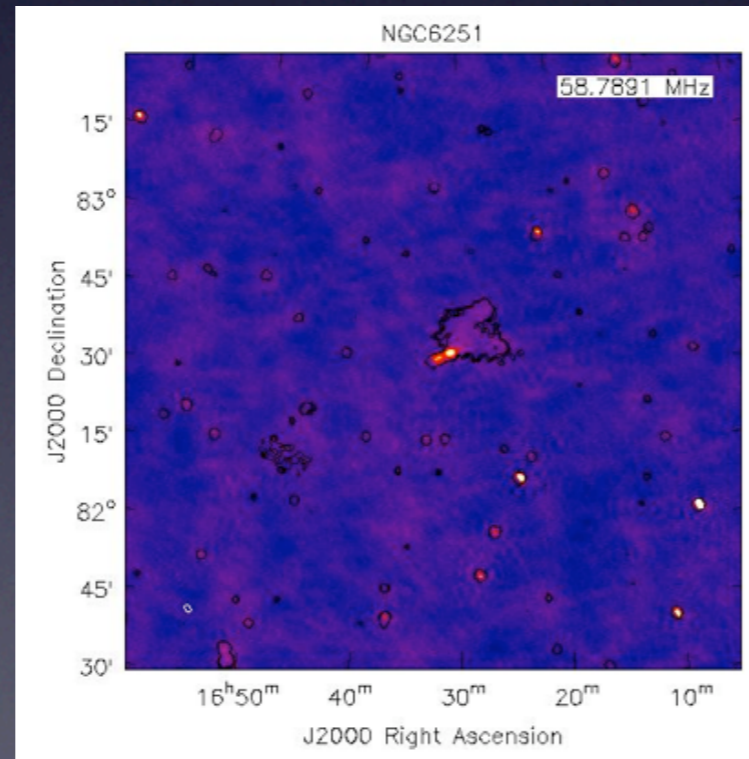
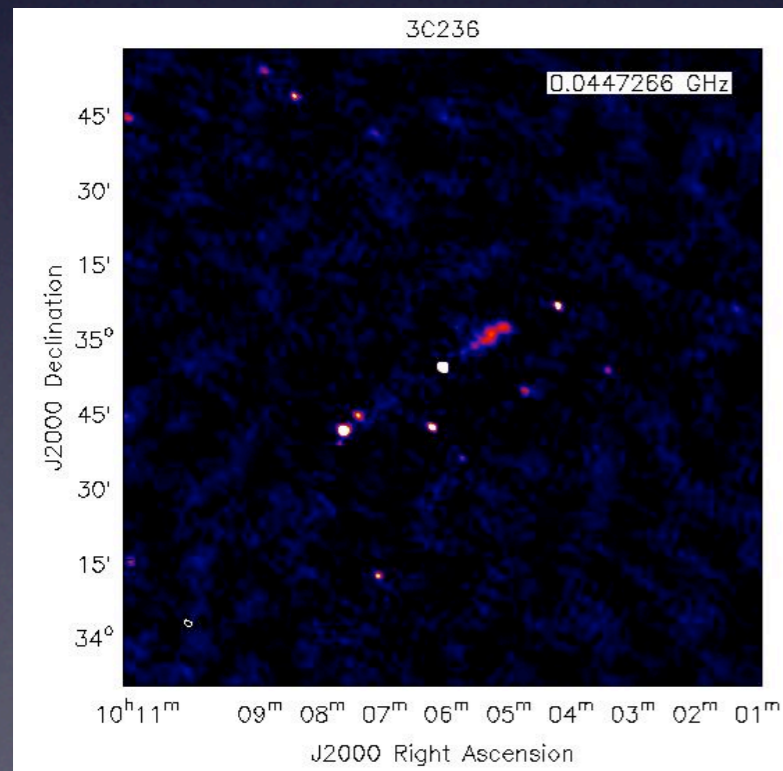
Emanuela Orru'

MKSP meeting Bologna

Collaborators: R. Pizzo, G. de Bruyn

Giant Radio Galaxies

- the goal: characterization of polarized emission in GRG to study AGN evolution in low density environments
- the group: G. De Bruyn (chair), E. Orru', R. Pizzo
- find calibrators in the Lofar frequency range
- the challenge: low surface brightness, difficult to model



LOFAR
LBA

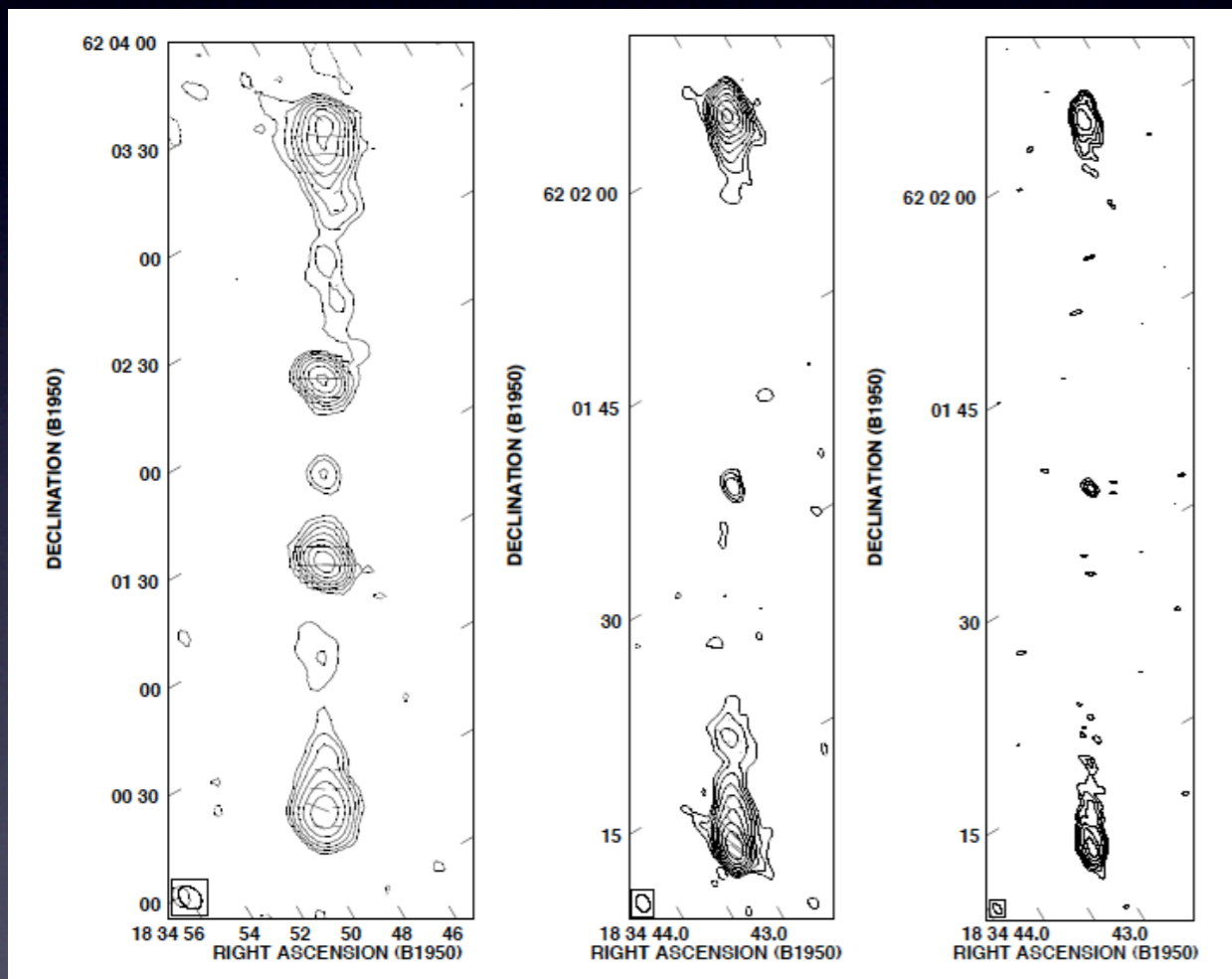
E. Orru'

courtesy of A. Shulevski

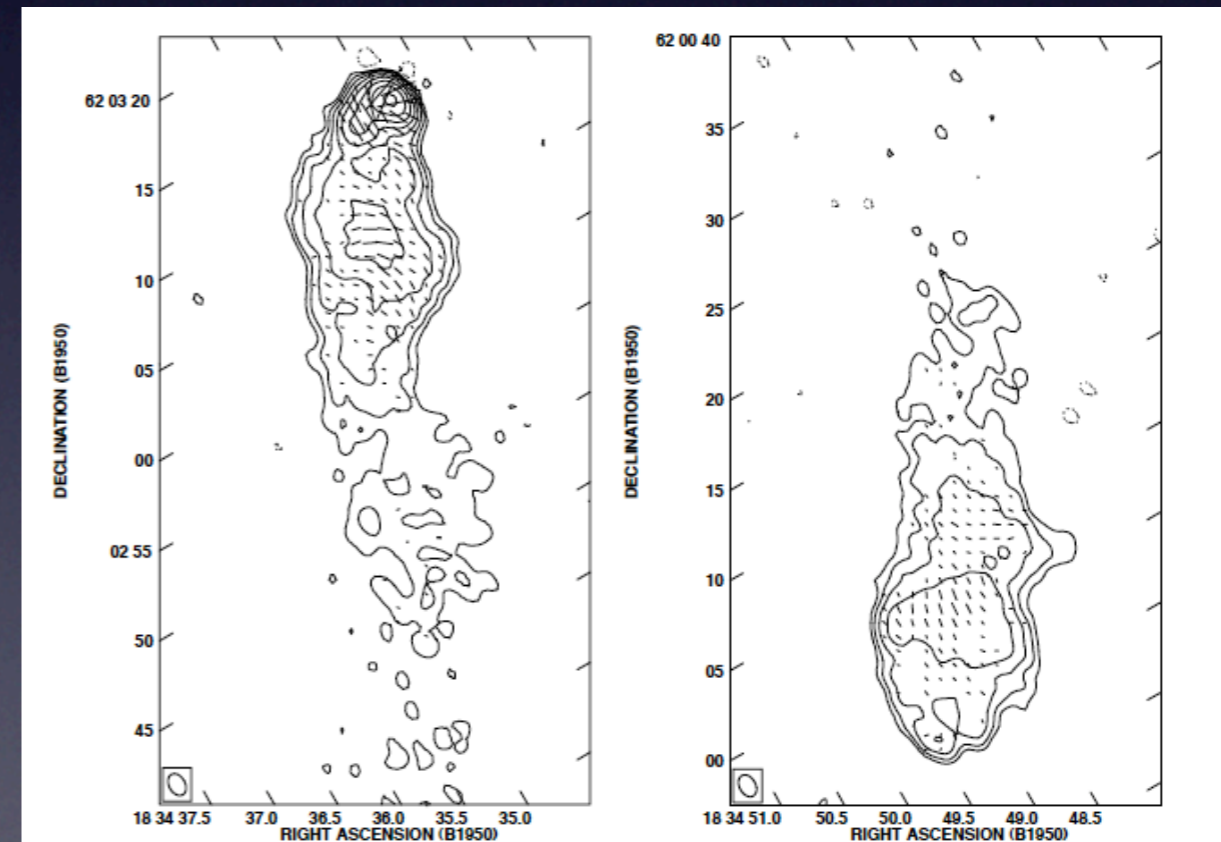
Commissioning targets

DoubleDoubleRG: B1834+620

- easy to model
- available WSRT model
- known to be polarized at 150 MHz (Ger priv. com.)

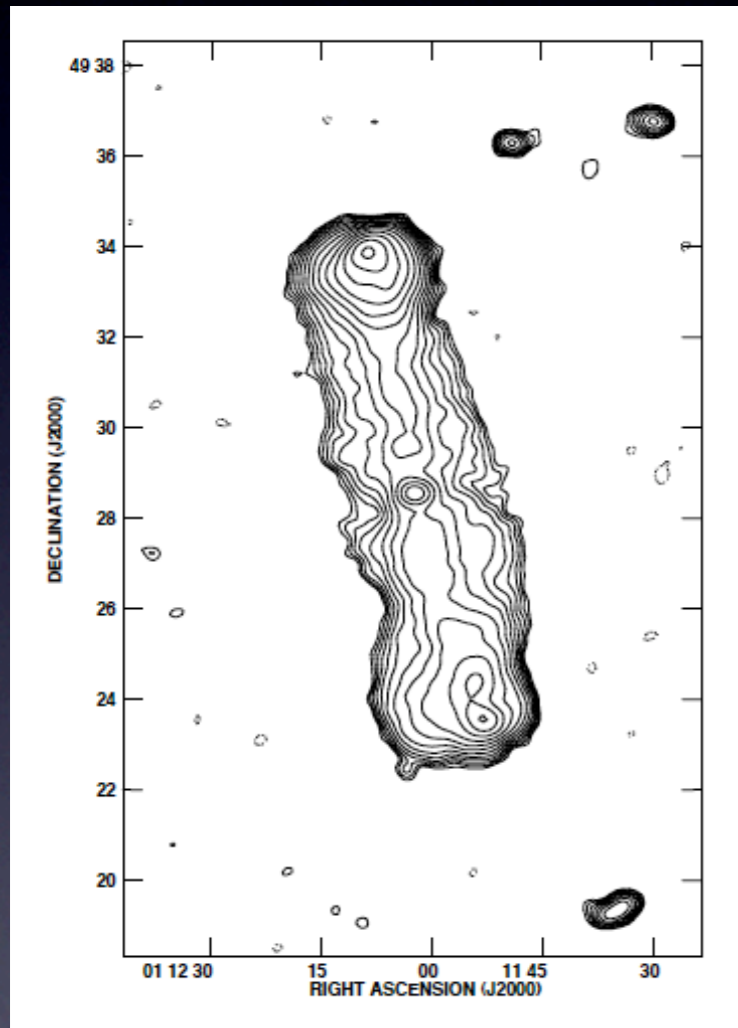


Schoenmakers et al. 2000

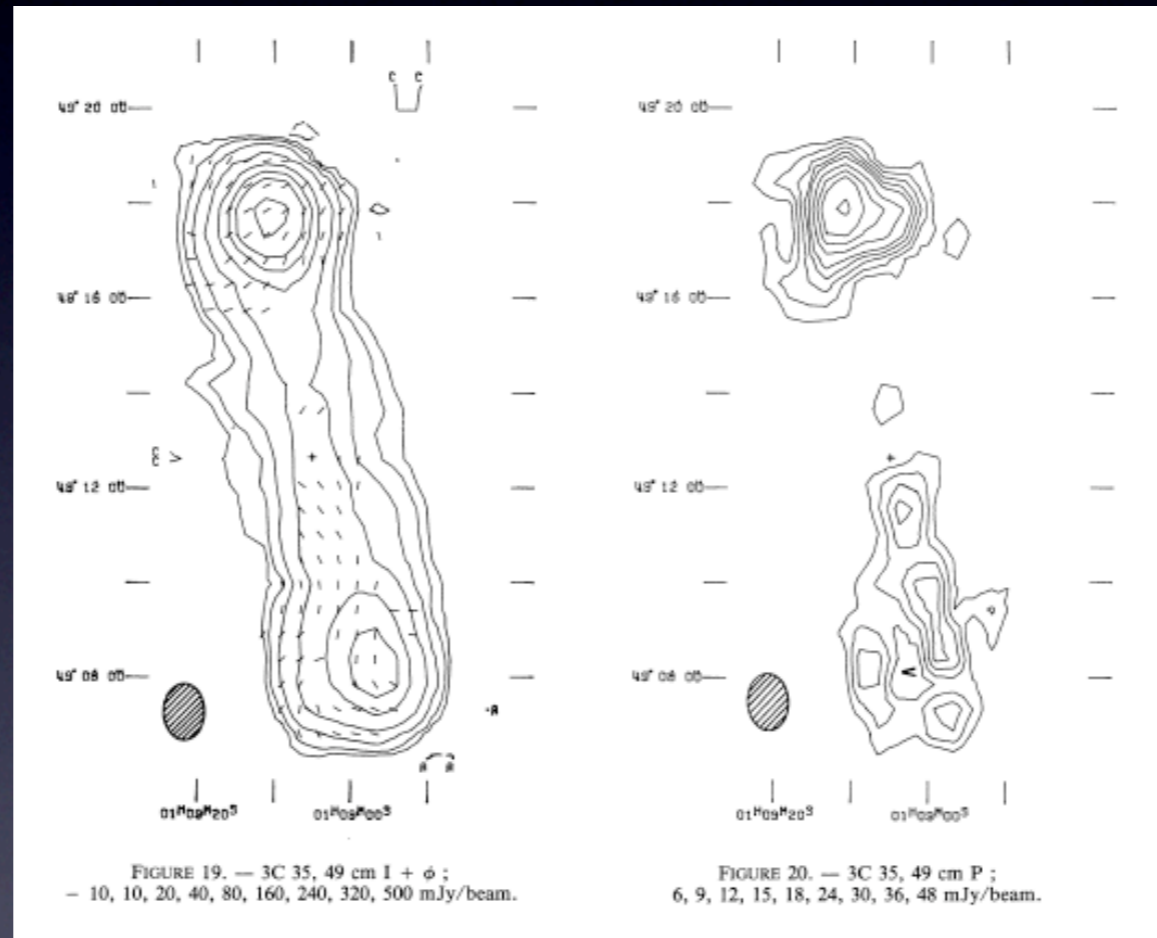


Commissioning targets

Giant Radio Galaxy: 3C35



Orru' et al. 2010

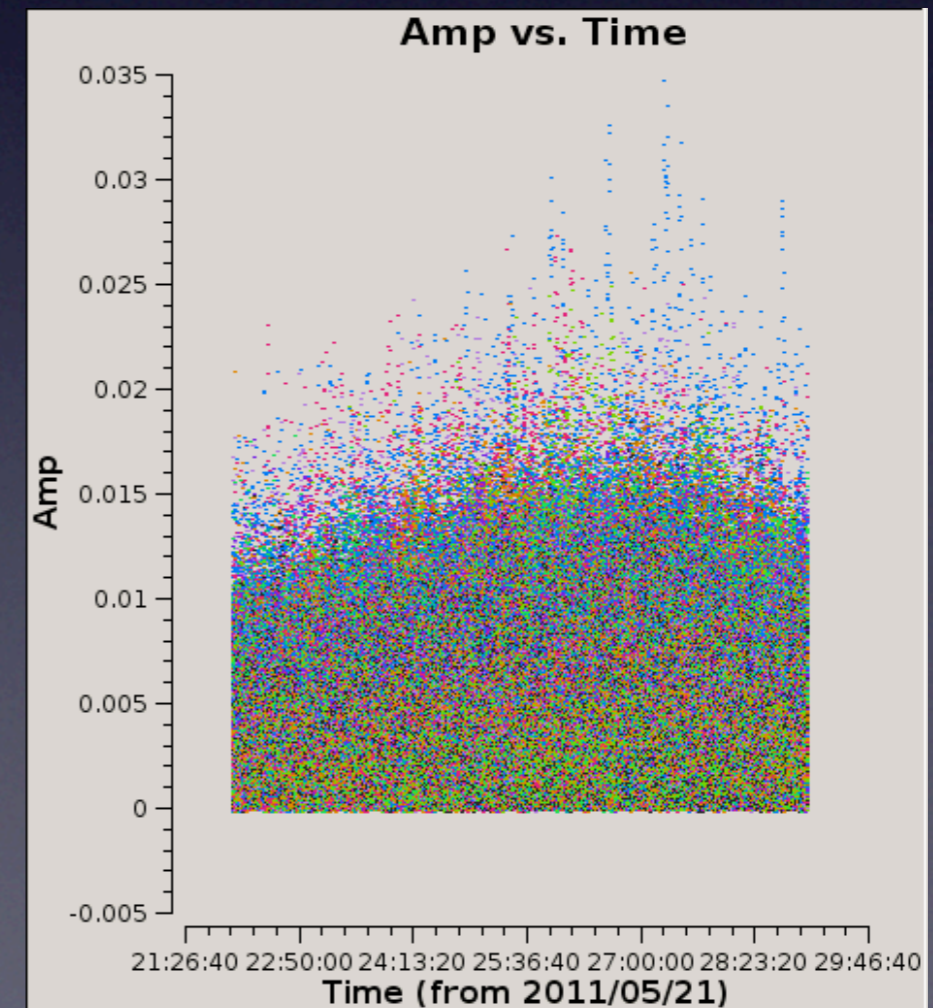
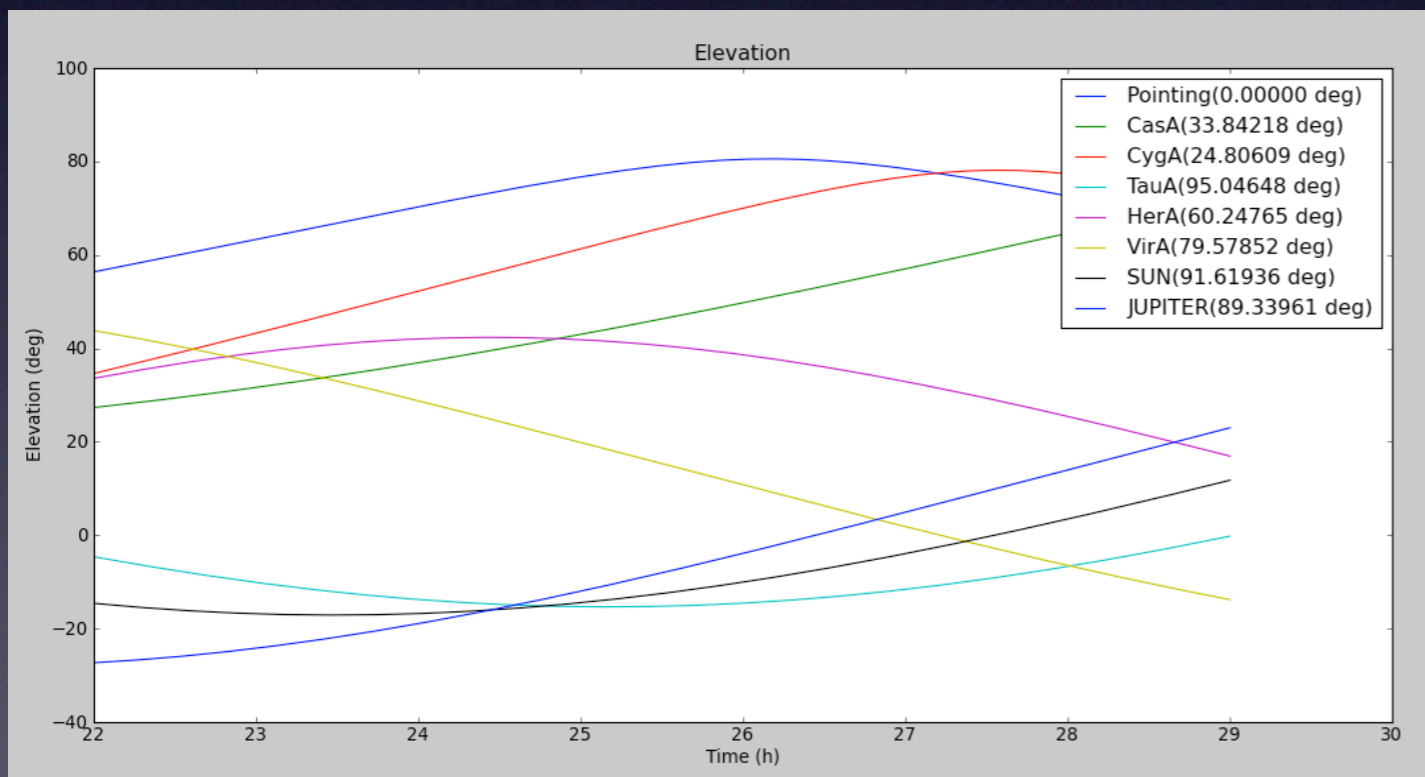
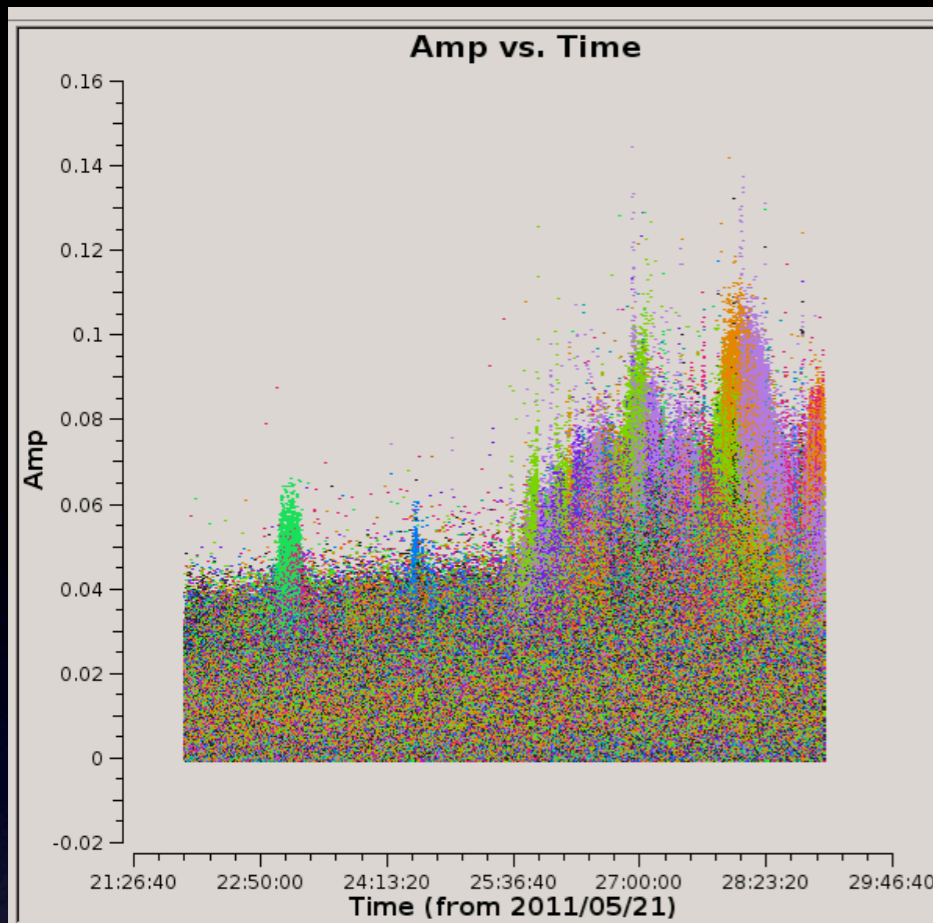


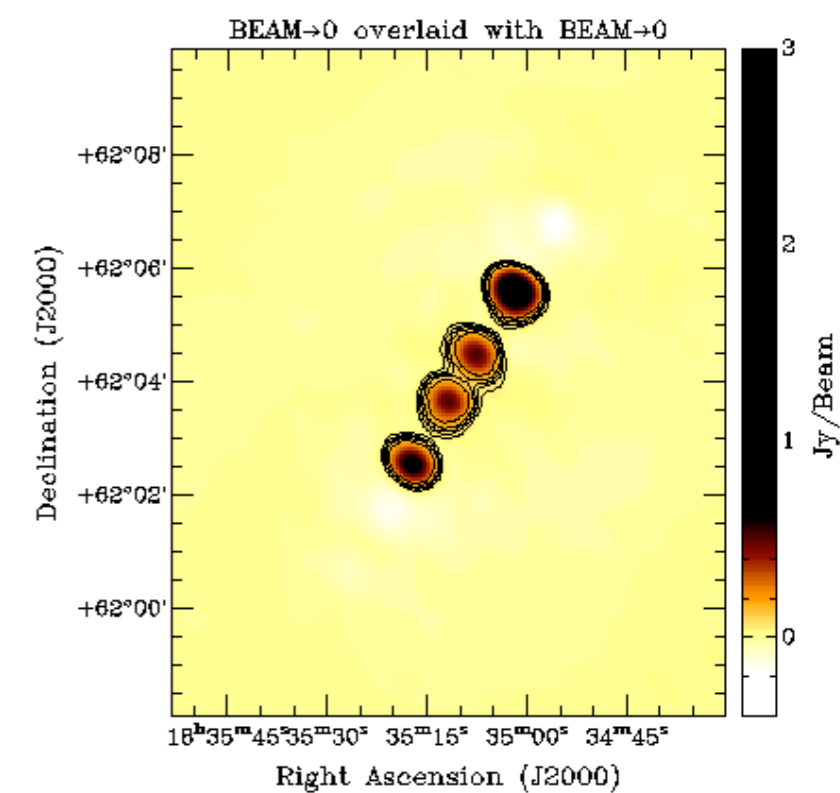
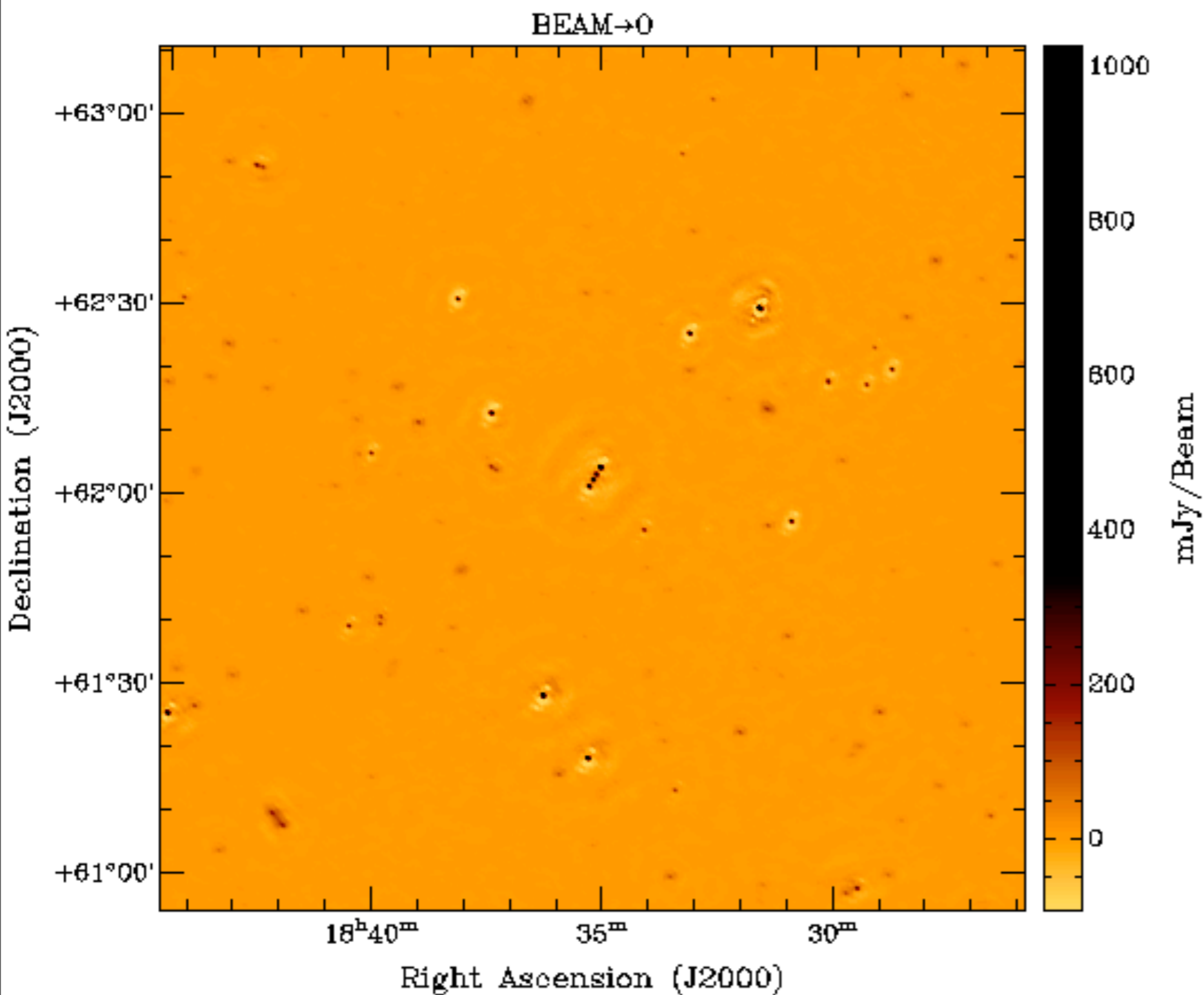
Van Breugel & Jaegers 1982

- known to be polarized up to 610 MHz
- Model VLA low freq. high resolution image

DOUBLE

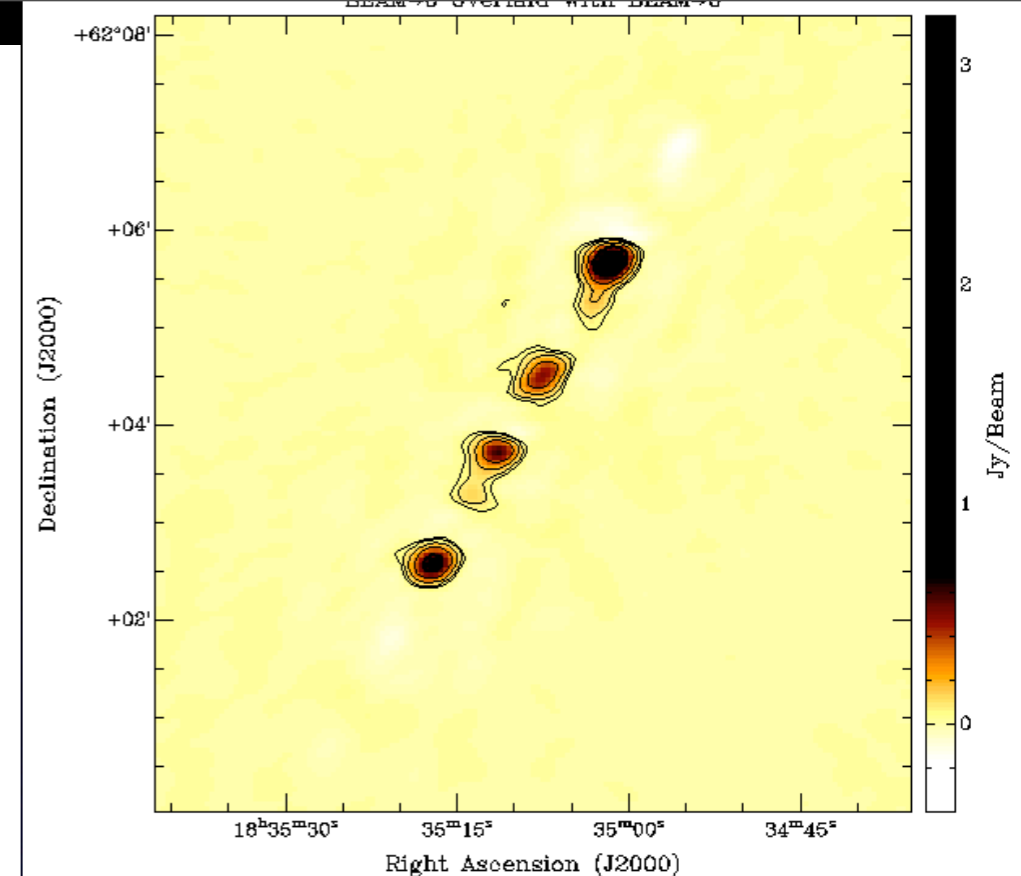
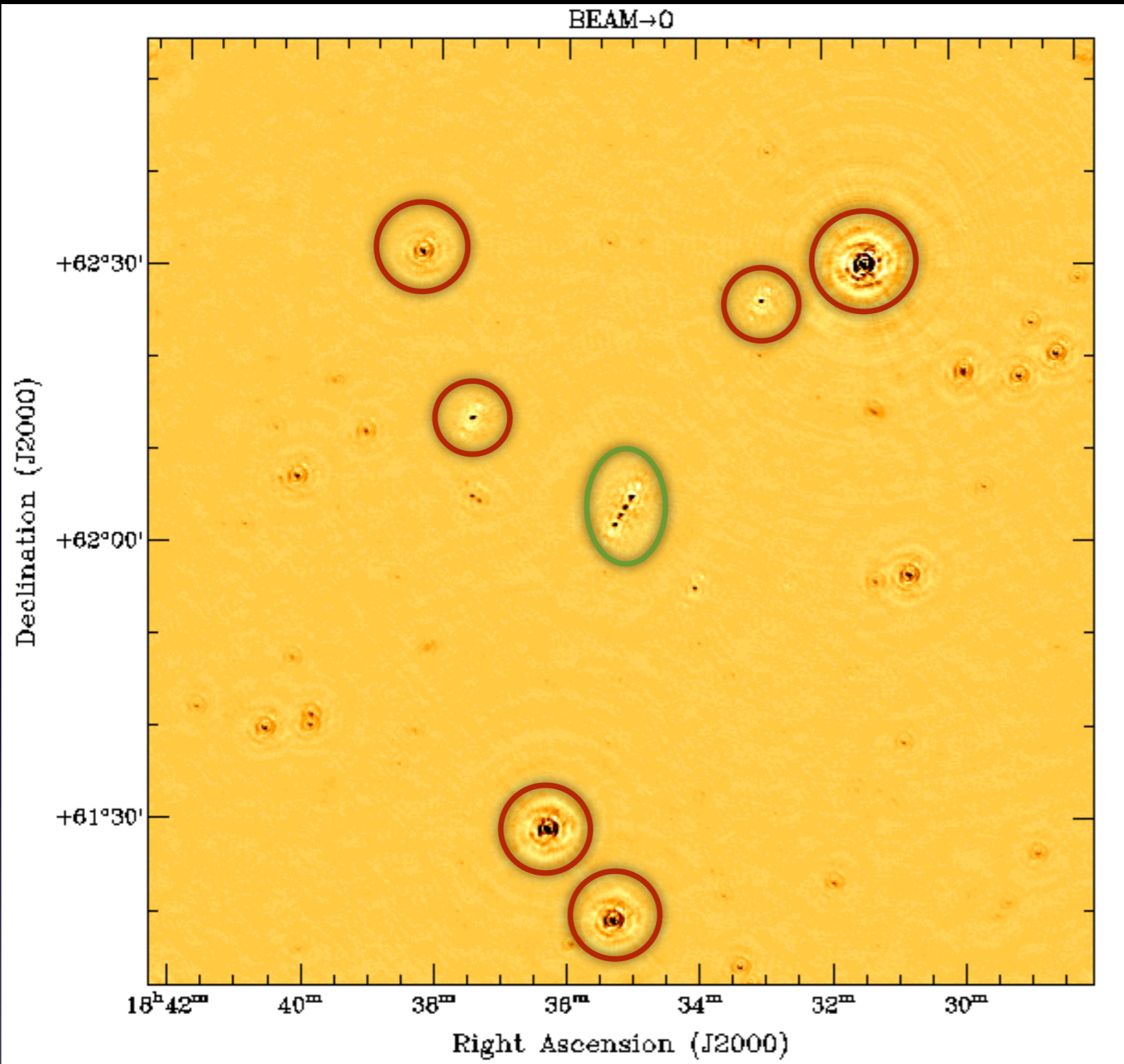
- 21-22 May-2011: 7h
- HBA freq. \sim 140 MHz
- 44 antennas IDE flagged + one with no data recorded
- 162 SB





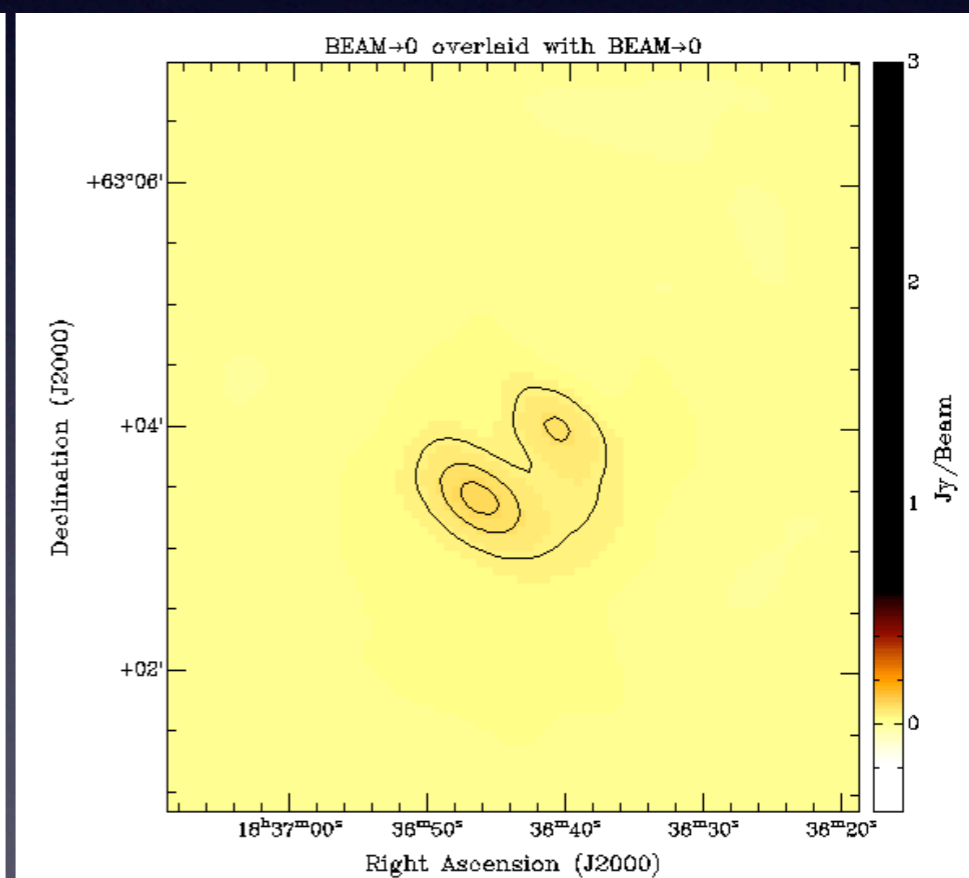
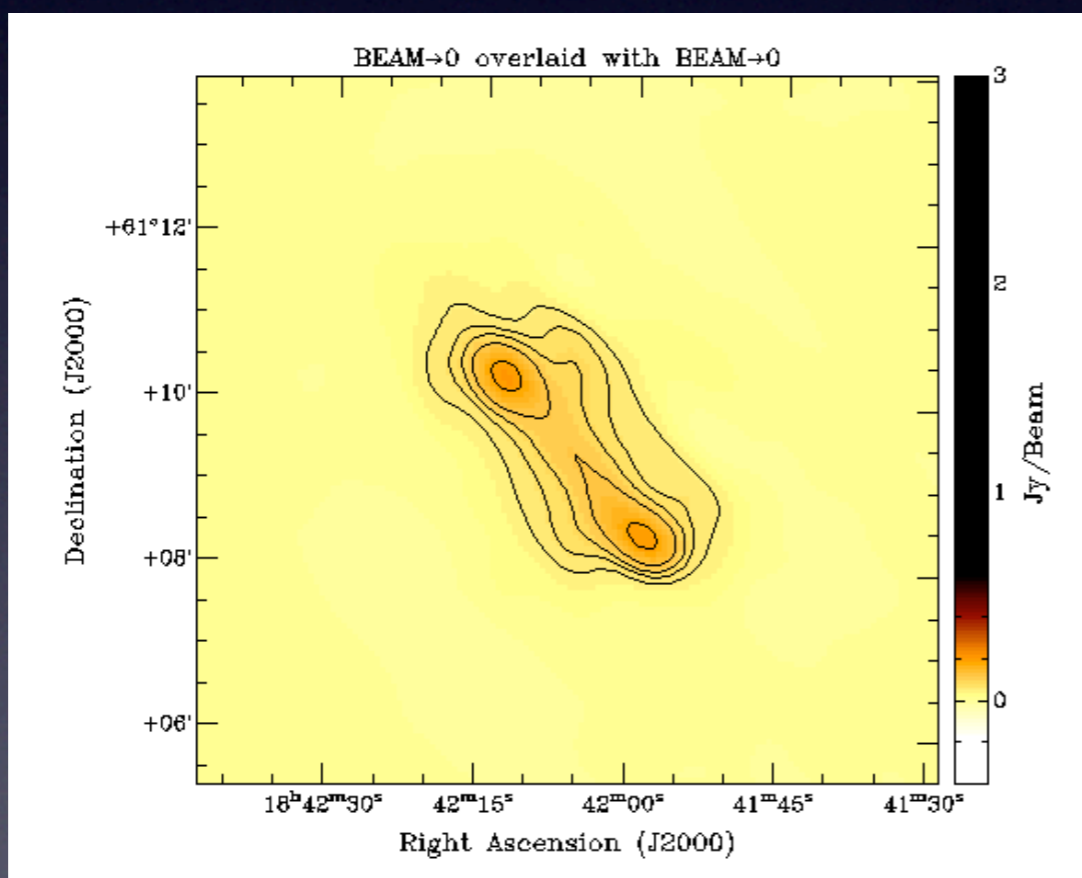
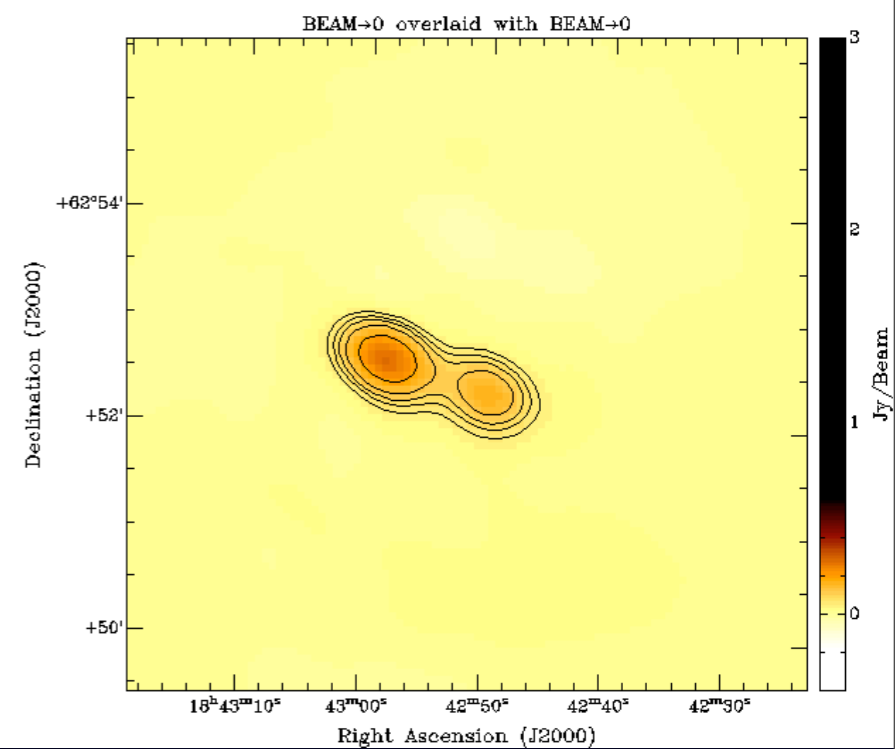
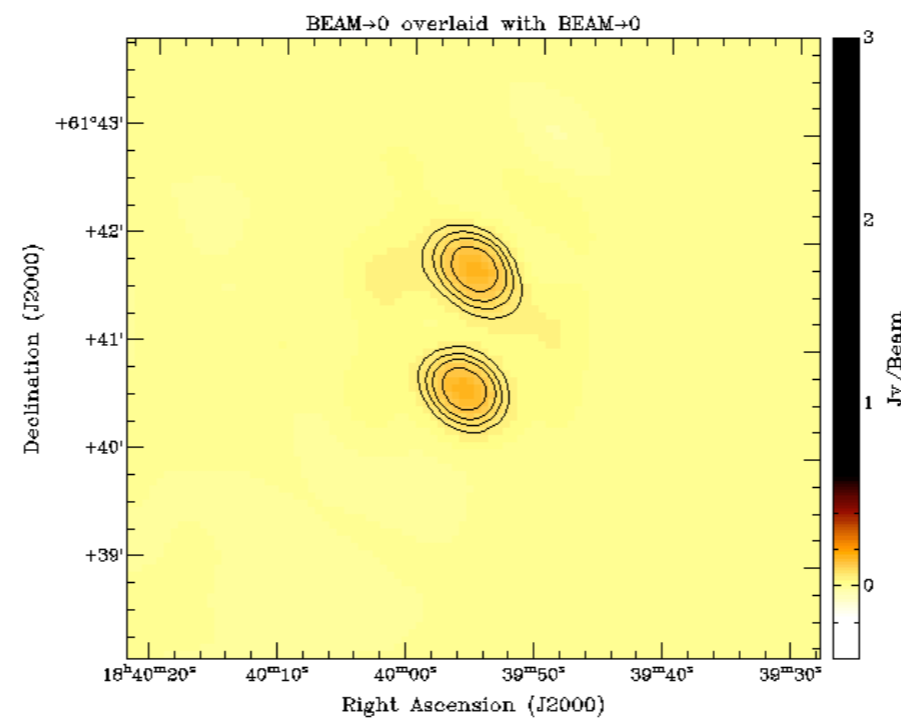
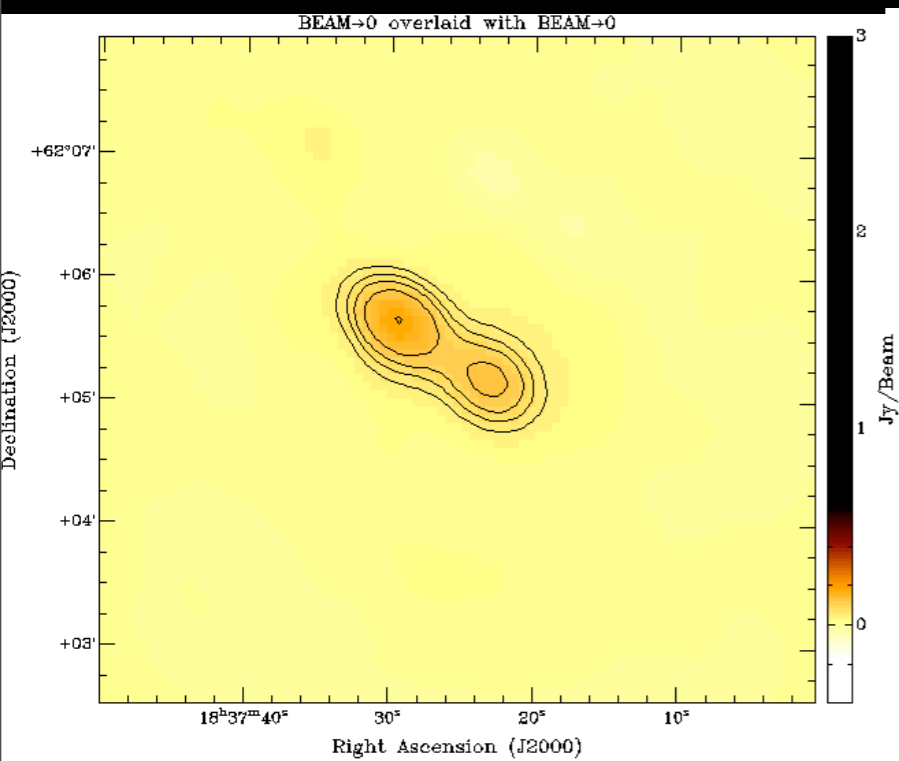
Low res=30"X25"
rms=2.5 mJy/beam

demix+BBS (SC)+flag+imaging

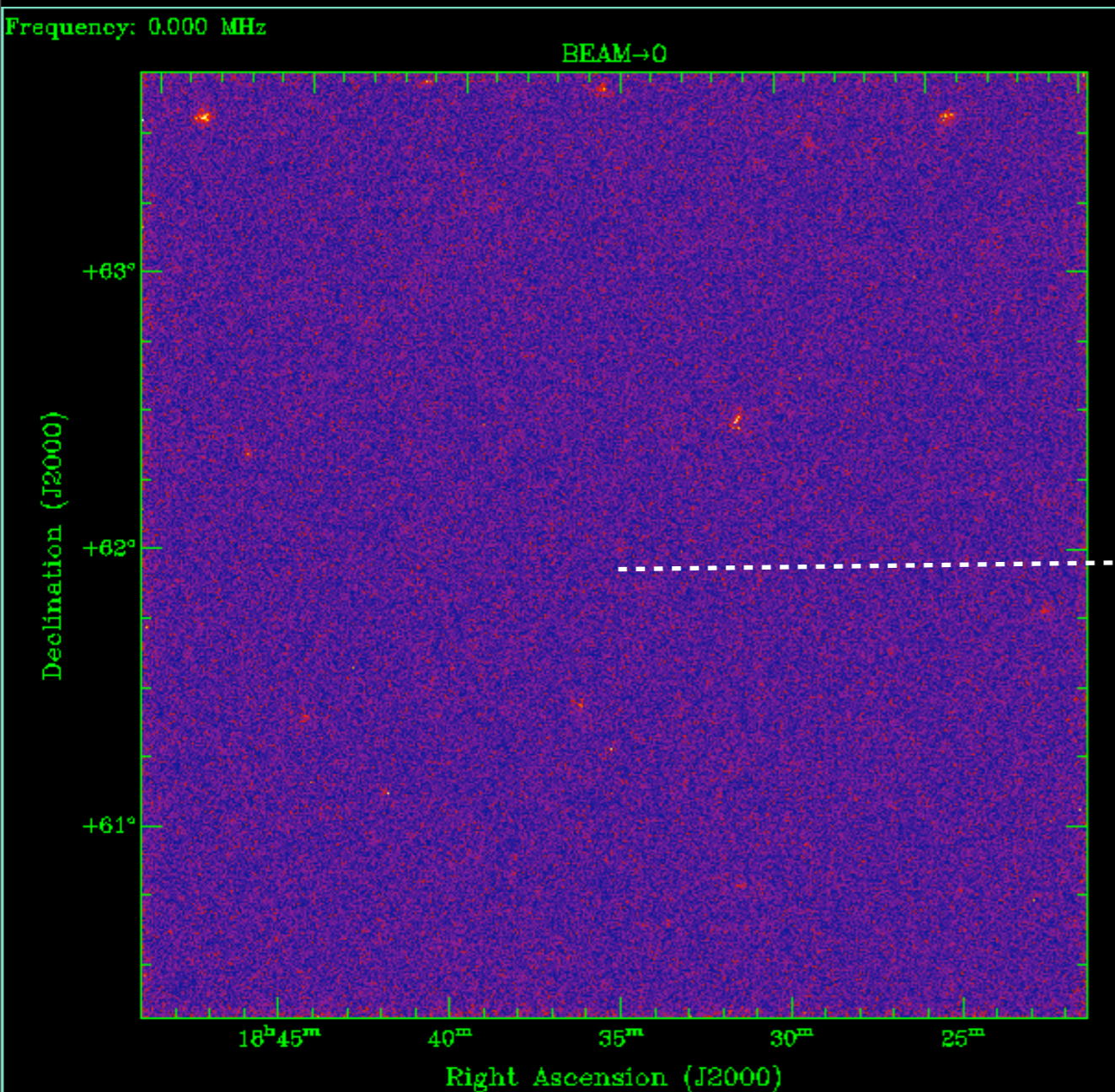


High res=22"X18"
rms=1.5 mJy/beam

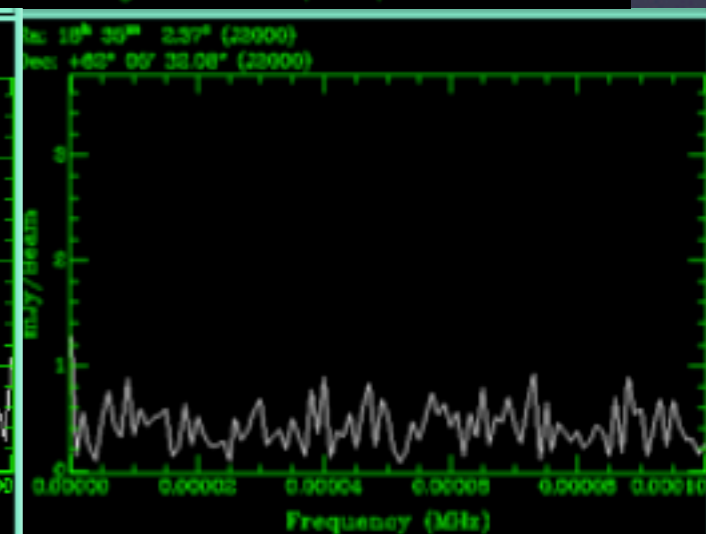
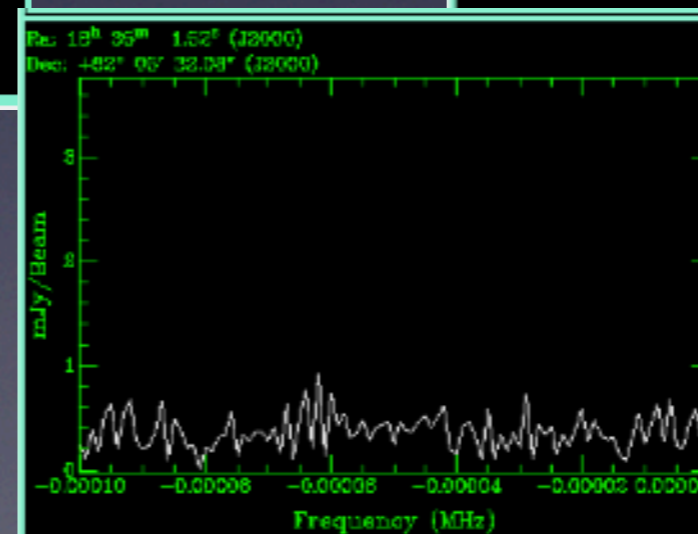
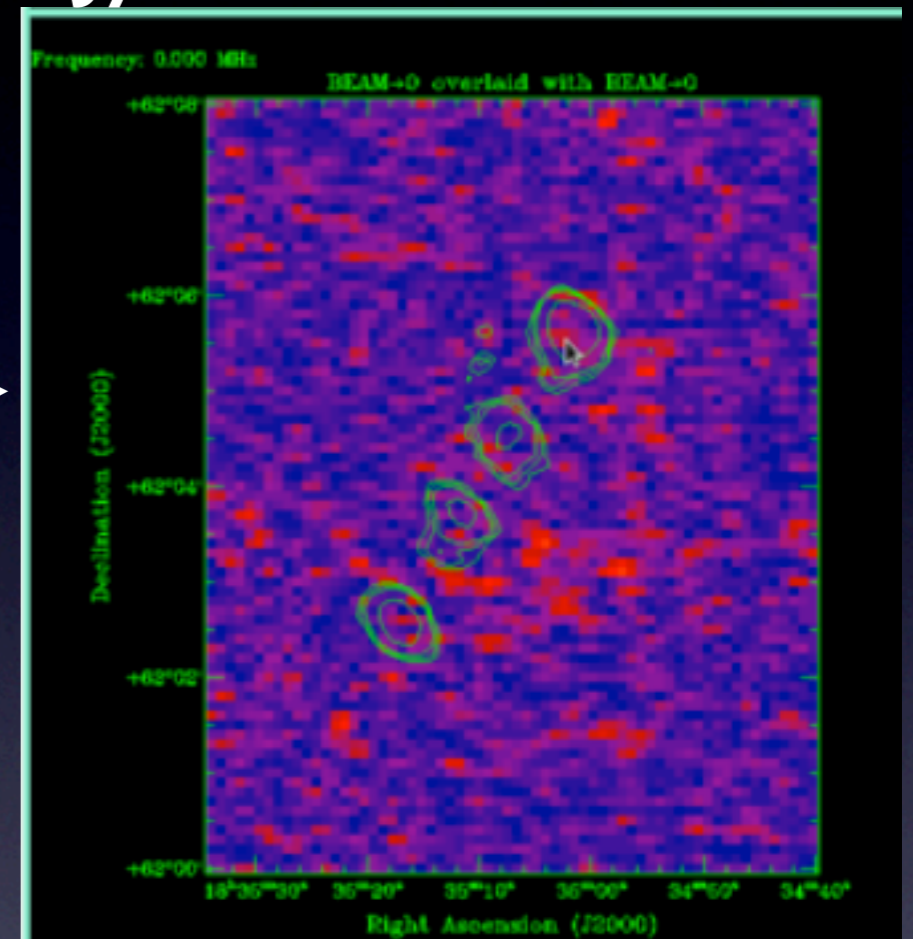
possible issues: model low res, DDE peak flux comparable,
variable beam/ionosphere? crash HBA dipoles



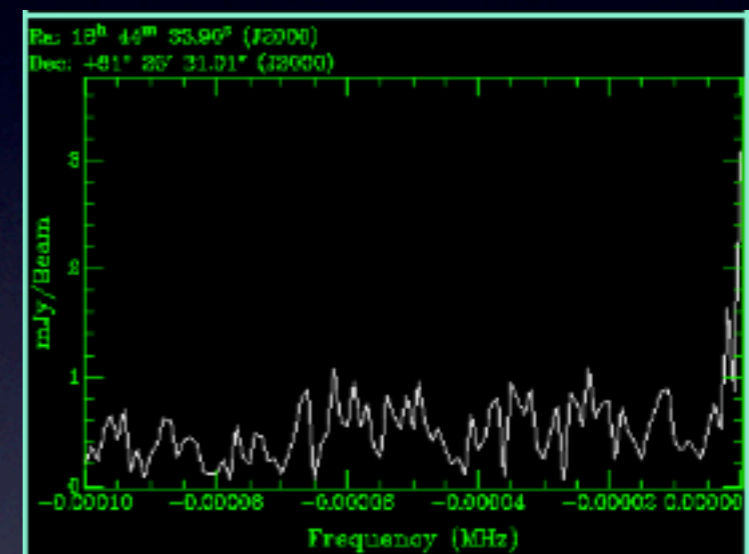
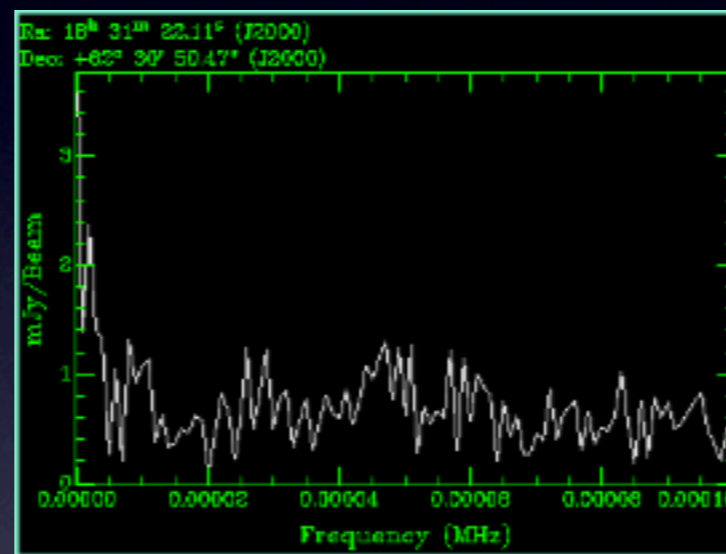
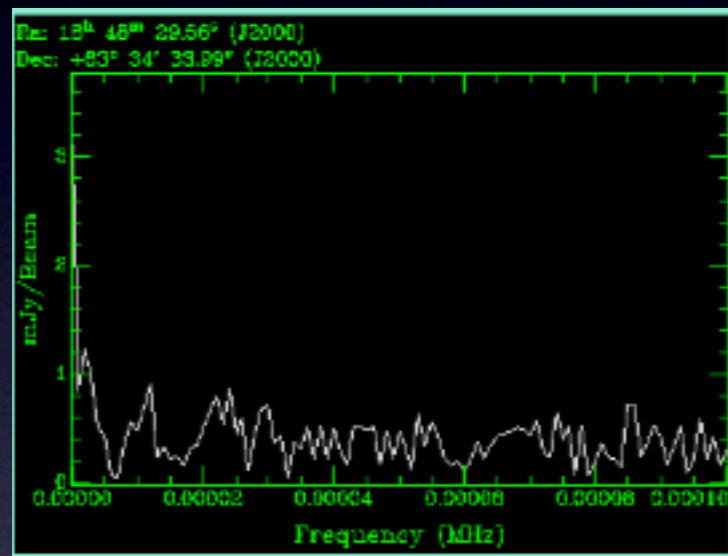
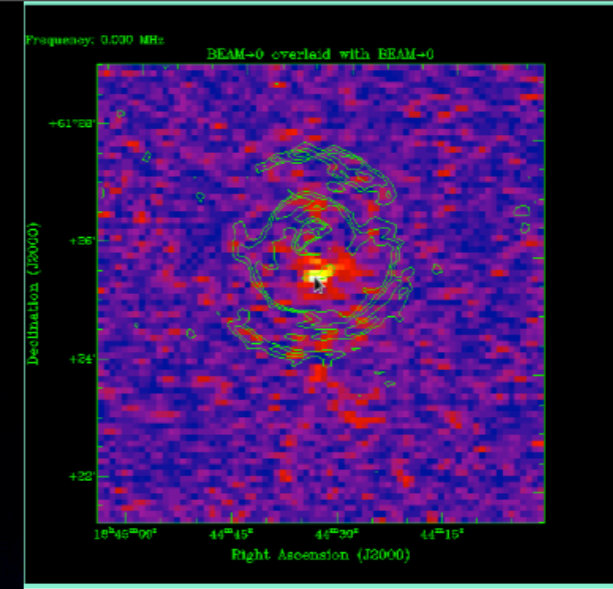
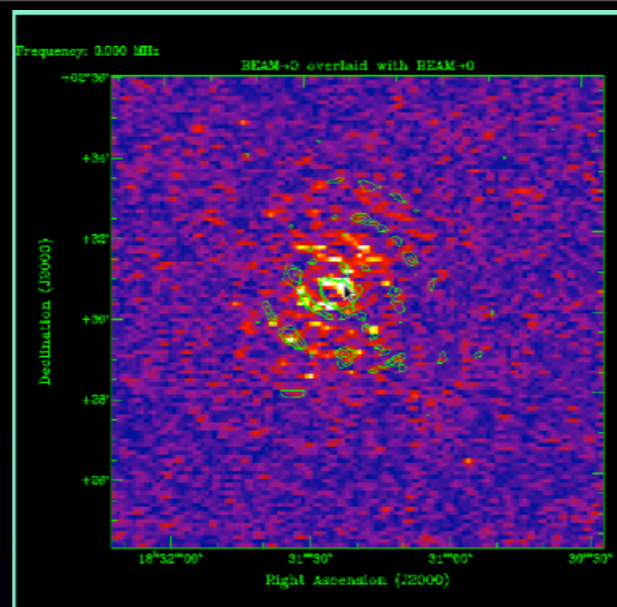
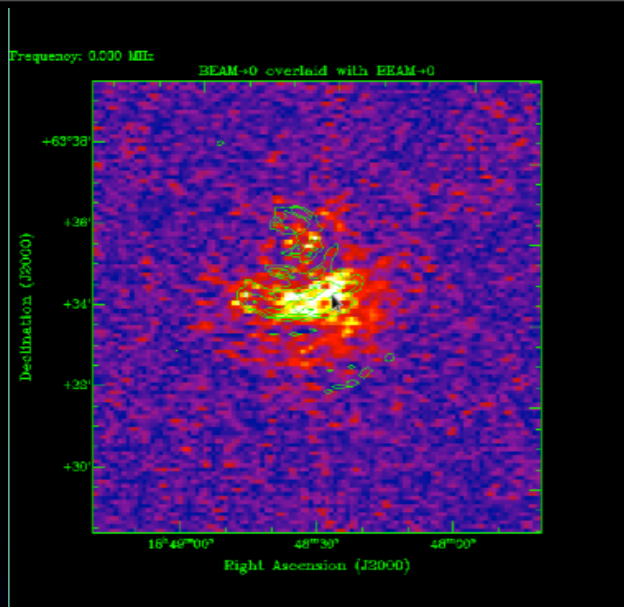
polarization: RM synthesis cubes



$\phi=0$
 $\text{rms}=0.3\text{mJy/beam}$



Trend instrumental polarization



Open issues: why no pol in double?

demix

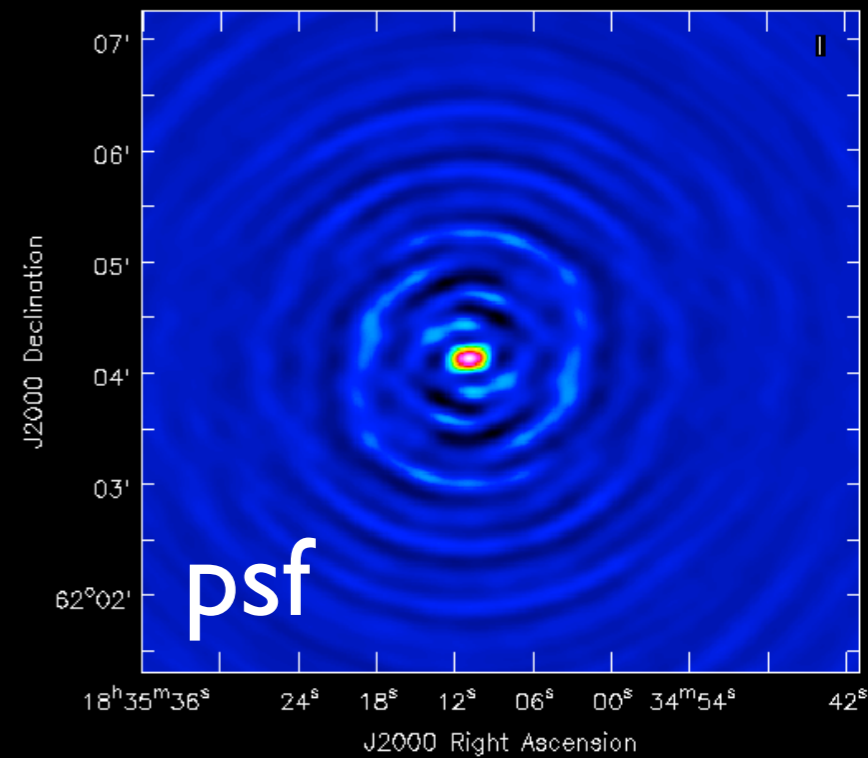
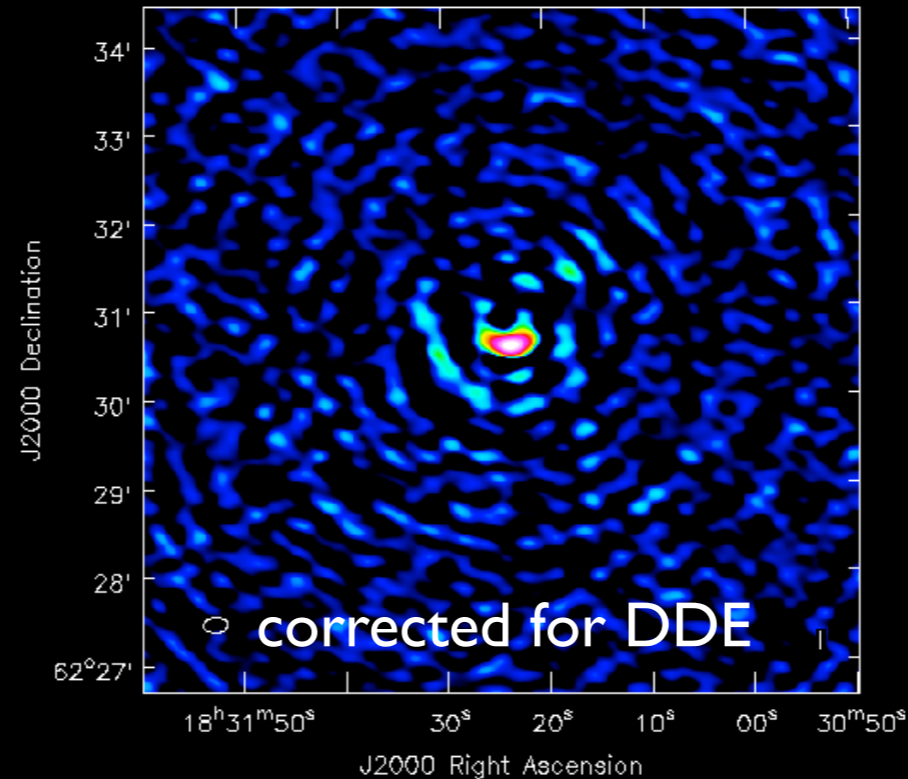
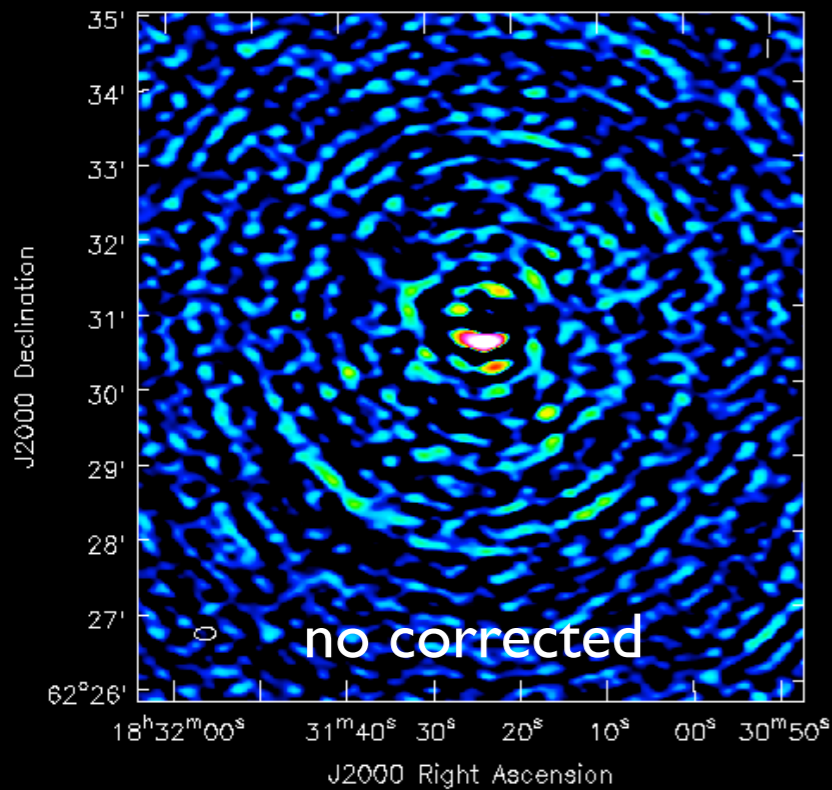
DDE artifacts

low S/N (162 SB out of 240)

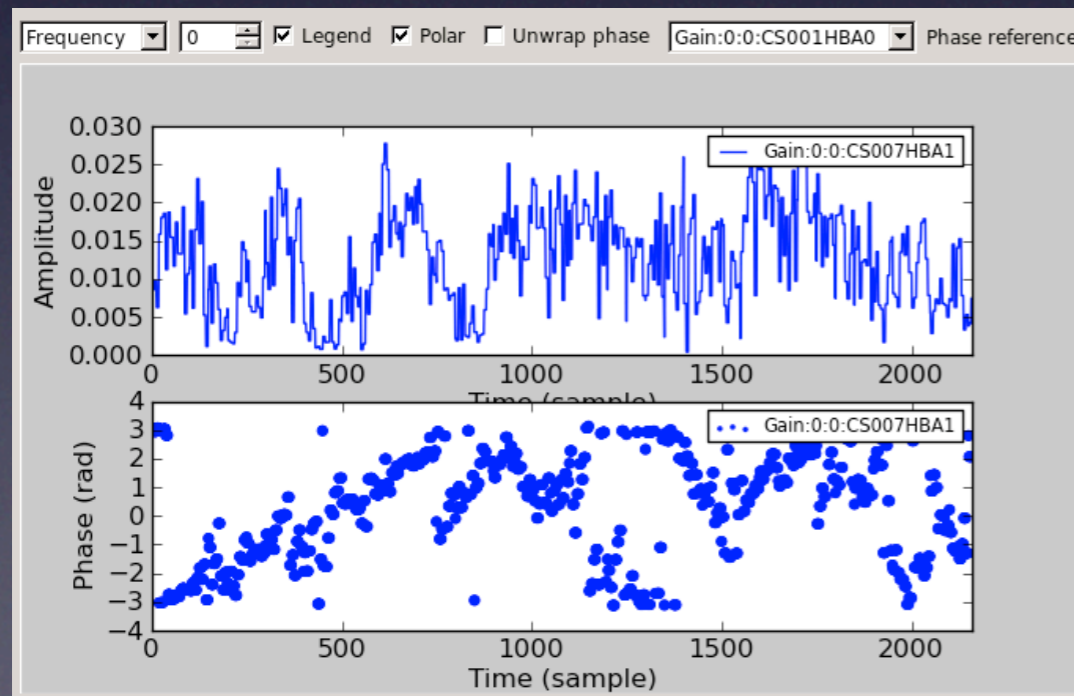
RM synth cubes analysis - emission in more pixels-lowering
resolution-beam depolarization

Possible Solutions and future tests?

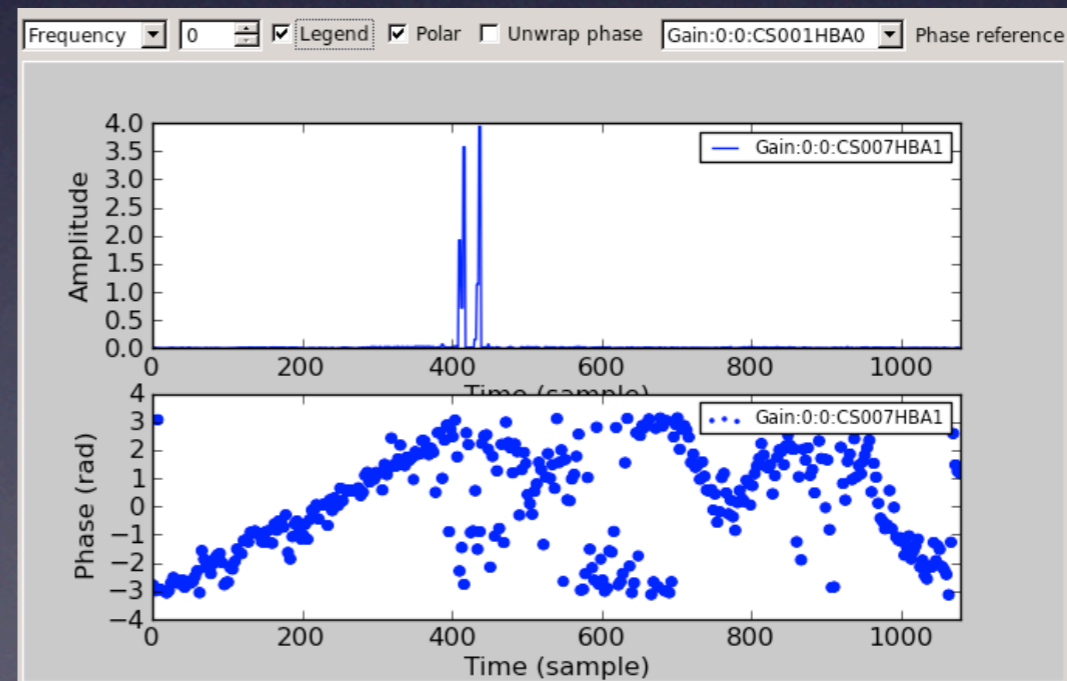
update the model



what if we do not use demix but subtract and calibrate in 2 step



2step



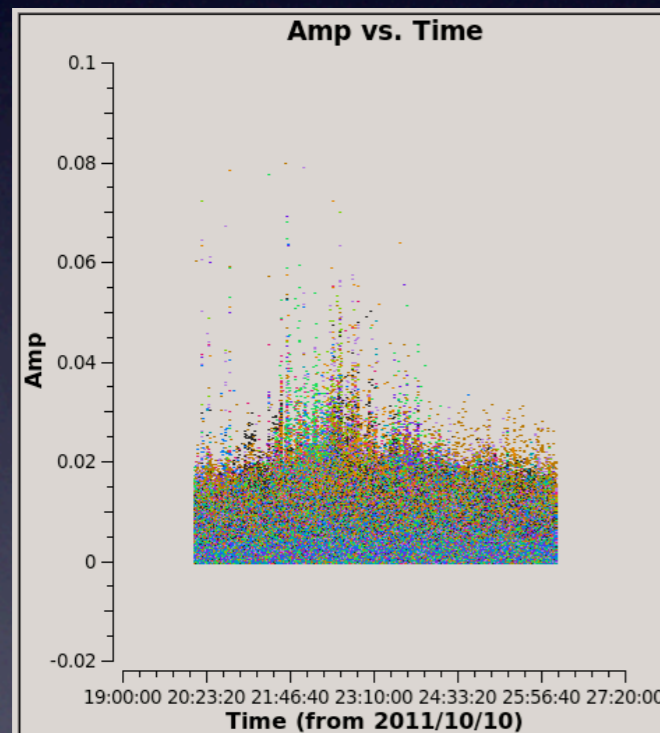
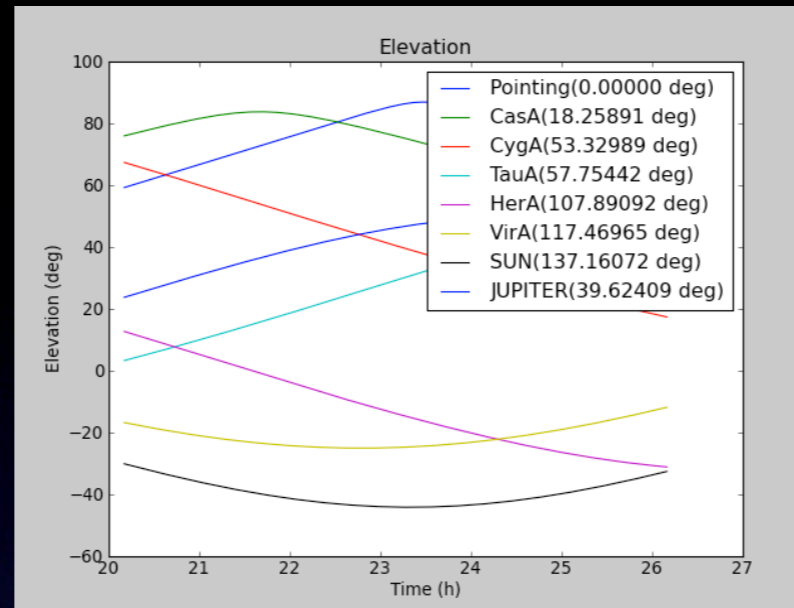
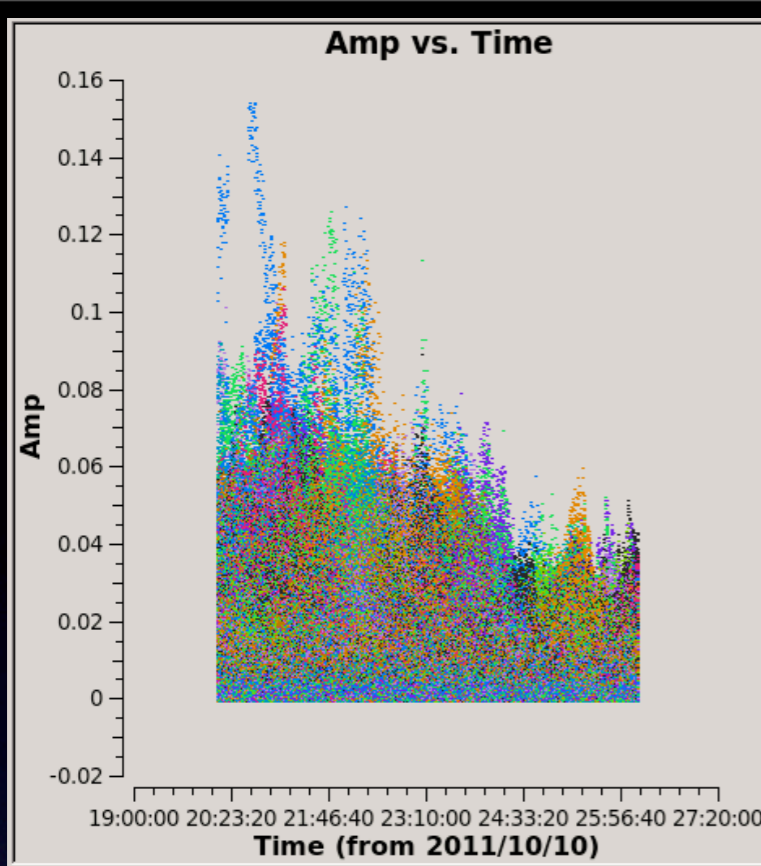
||

demix

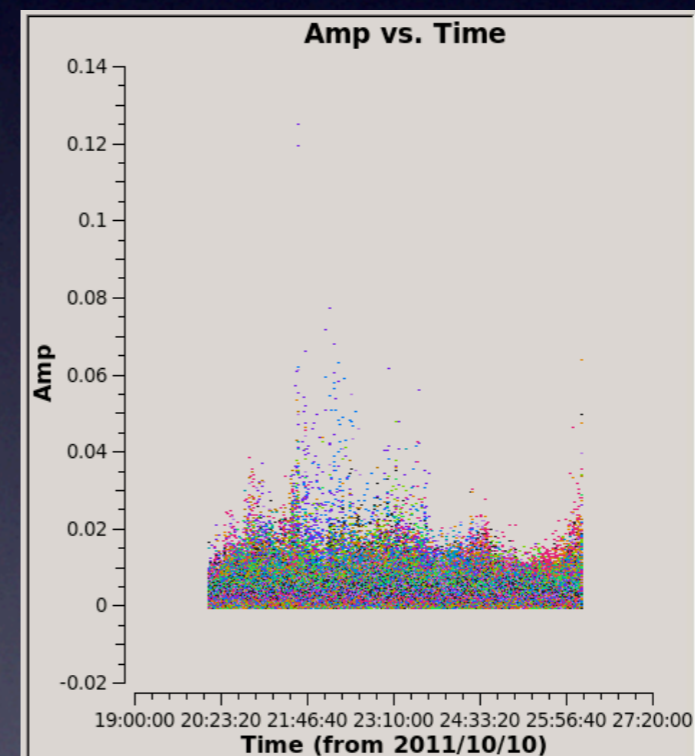
3C35

A-team

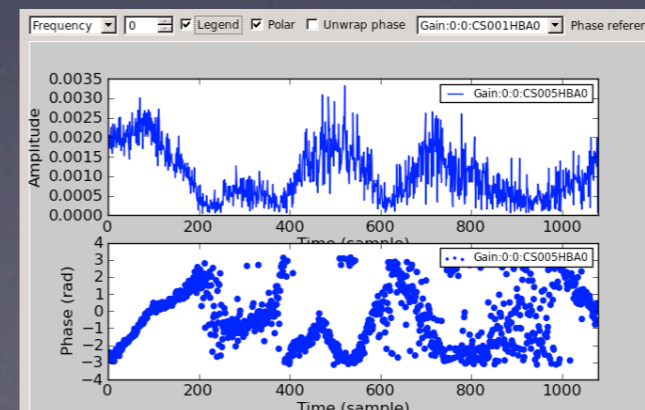
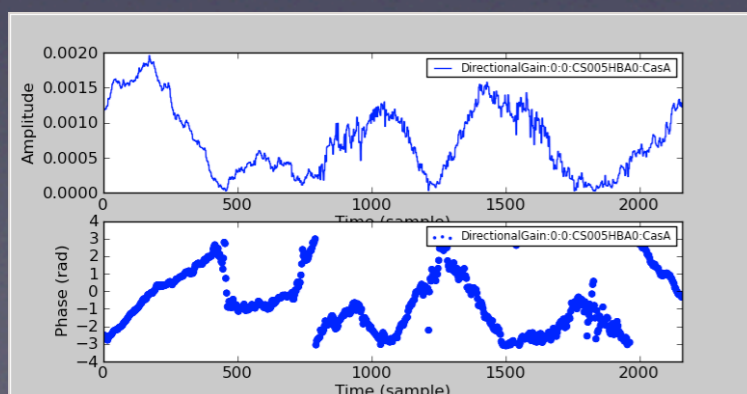
- 10 Oct-2011:6h
- HBA freq. ~ 140 MHz
- 50 antennas



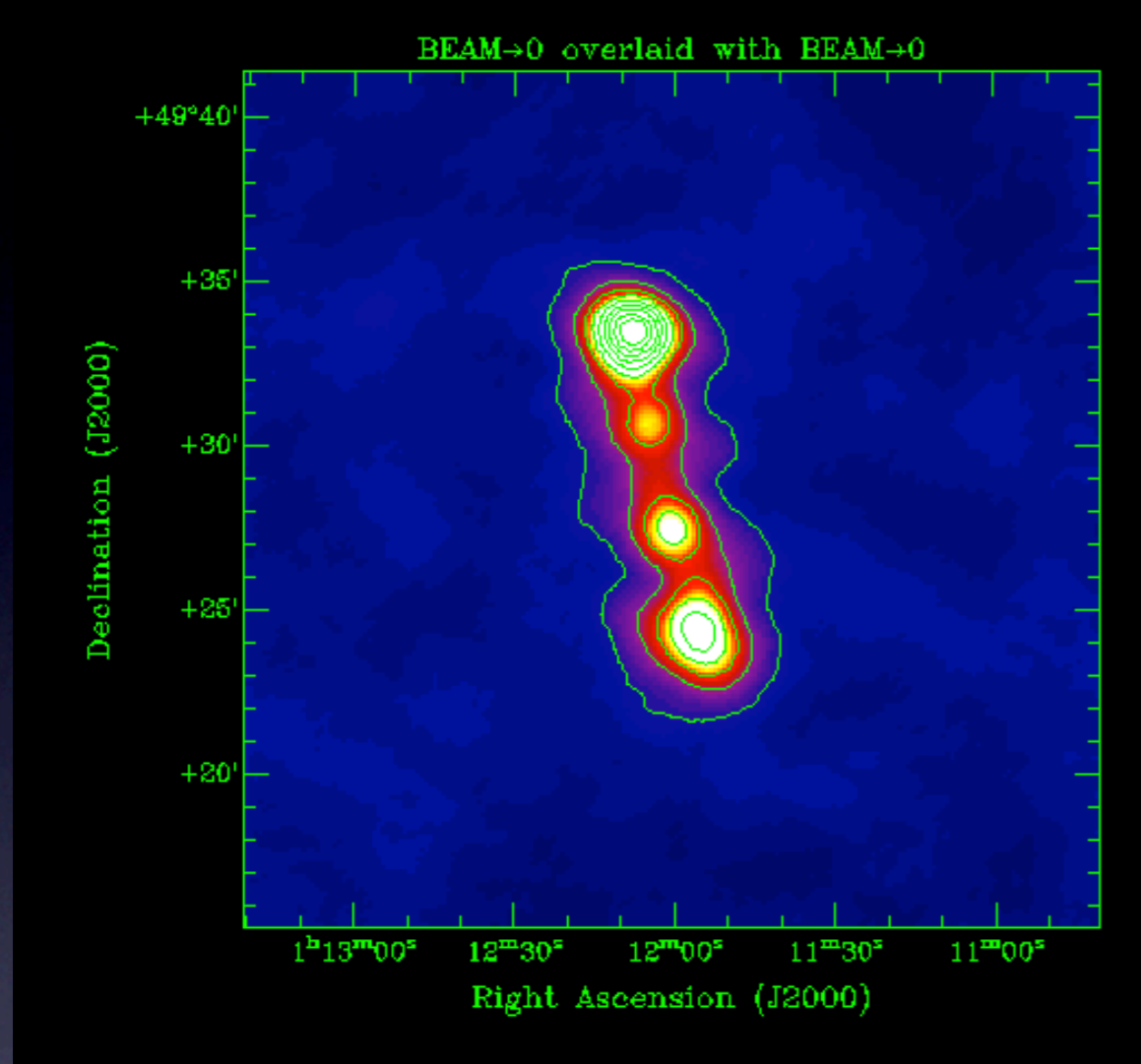
Subtract DDE



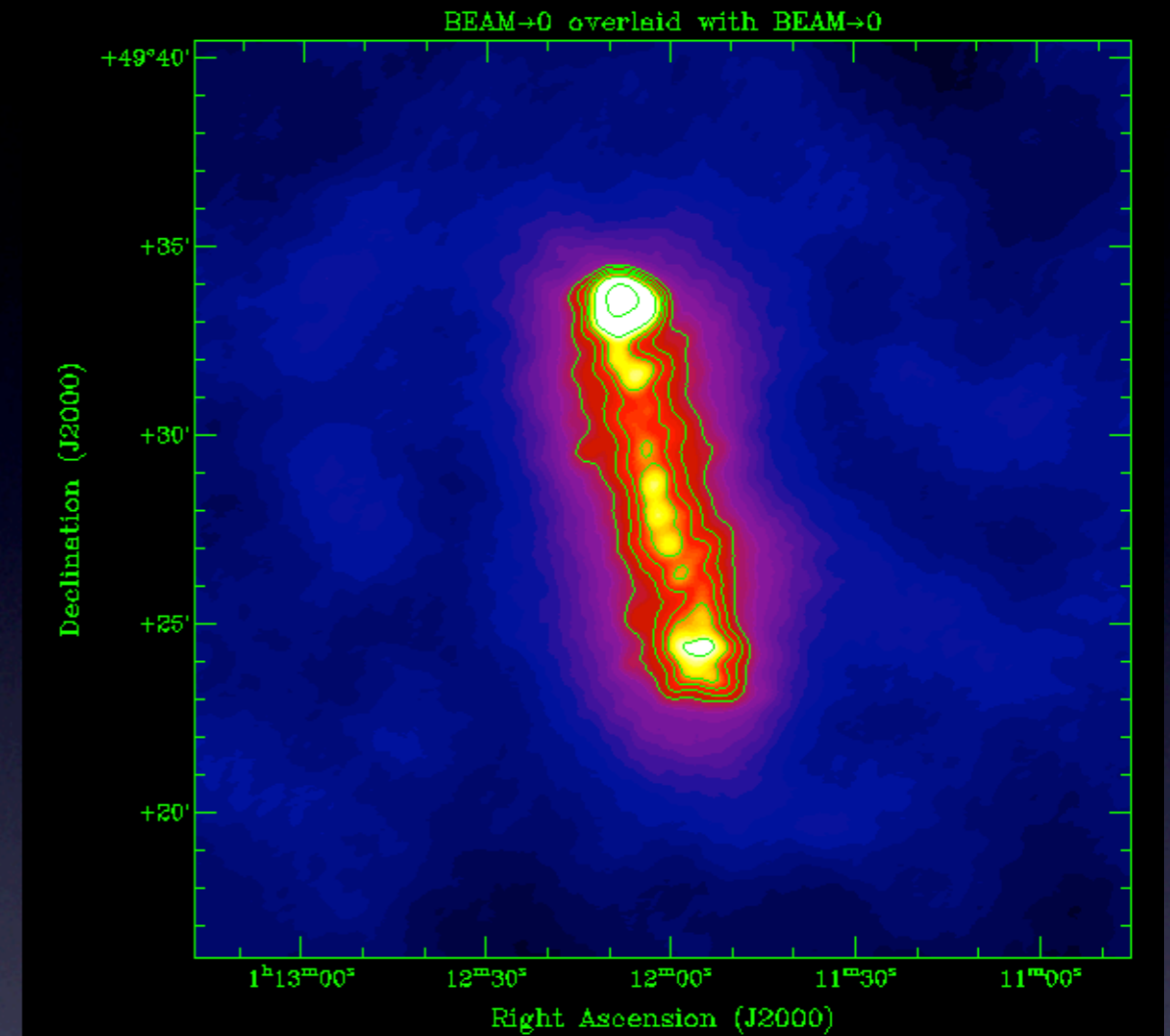
Demix



preliminary images



res=99"X90"
rms=35 mJy/beam demix

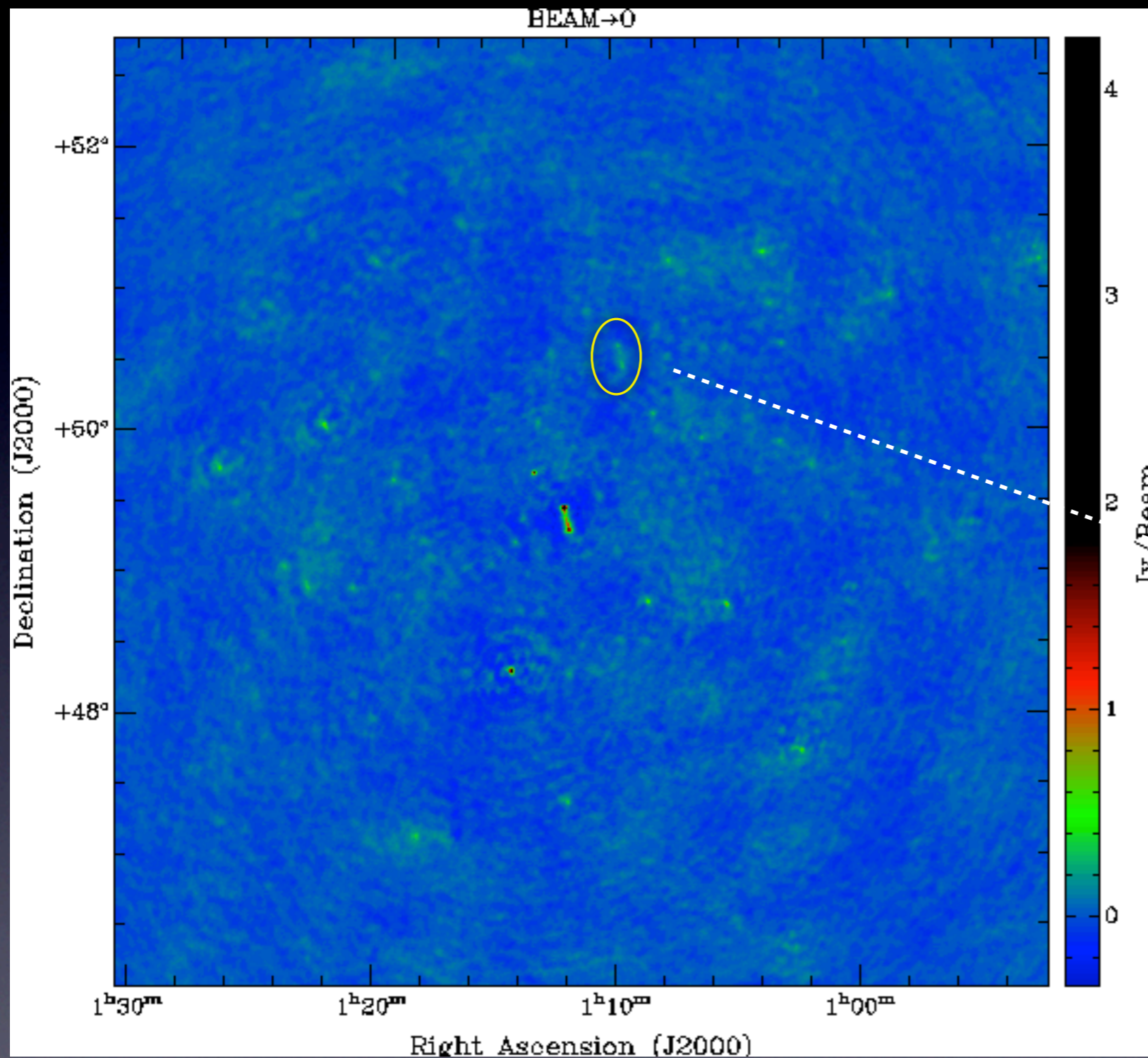


res=45"X39"
rms=30mJy/beam demix

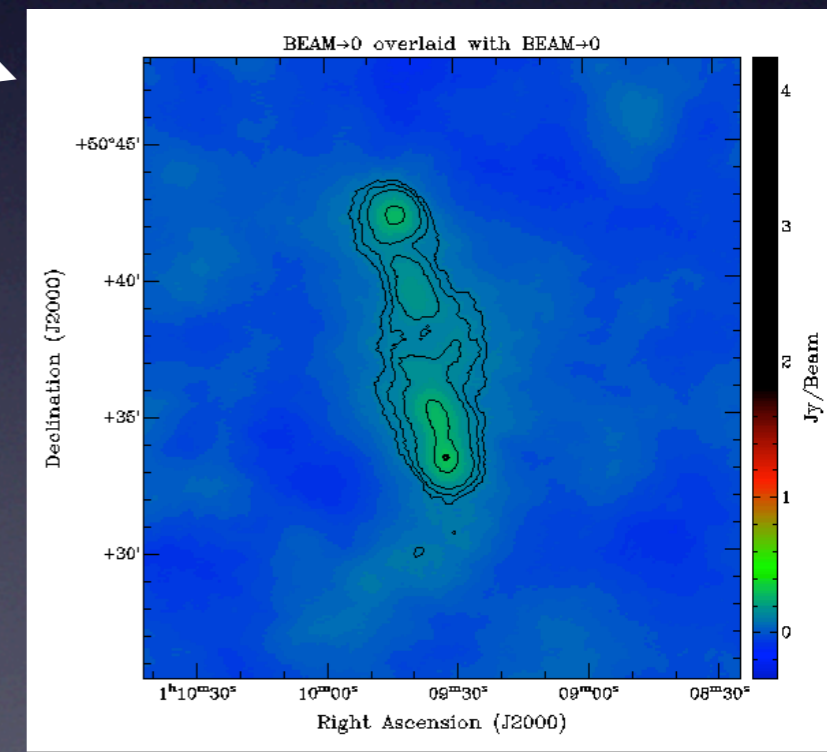
low level RFI after A-team subtraction
pattern short BL
Q UV noisy no evidence of polarized emission

future: model combined
my thesis wenss
point and shapelets
flag low level RFI

preliminary images



res=77"X70"
rms=30 mJy/beam
2step method



Summary

- B1834+620: total intensity and P images. Artifacts offset sources. Instrumental polarization.
- update the model at high resolution, investigate on the reason for SB crashes
- 3C35 preliminary calibration. Total intensity show a pattern, no polarized emission observed in Q U V.
- produce a combined model with point-shapelets, test demix vs 2 step subtraction, residual RFIs after A-team subtraction.

Limits

- calibration DDG correct in several directions - computing capacity -single vs global
- subtraction method demix or 2step subtract
- RM synthesis slow, memory limited, limited in pixels (high resolution small FOV) and Phi range and shift R.A.
- RM cube visualization limits @ 2Gb (kvis).